

## **Polarized Neutrons**

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Oxford School on Neutron Scattering - 2024





Polarized neutrons can be used to enhance almost any neutron scattering experiment, either by

- 1. Separating and providing additional information on the **sources of scattering** (nuclear, magnetic, incoherent)
- 2. Improving the **resolution**





### Outline

### Definitions

Neutron spin, magnetic moment, precession Polarization, flipping ratio, polarizers, flippers

### **Polarized neutron scattering**

Polarized neutrons interaction with matter

Separation of sources of scattering

### **Neutron spin-echo**

Principle of NSE - using the neutron spin as measuring tool







### Definitions: Neutron magnetic moment

Particles with spin are experimentally observed to possess a magnetic moment

 $ec{\mu}=\gammaec{s}$  where,  $ec{s}\equiv\pm s_z=\pm 1/2\hbar$  and  $\gamma=-183$  MHz / T for neutrons



Quarks also have s =  $\frac{1}{2}$  quantum number, therefore if the up and down quarks are antiparallel, the neutron also has s =  $\frac{1}{2}$ 

The magnetic moment calculation is tricky – but can be measured accurately

Often we see the neutron magnetic moment given in terms of **nuclear magnetons** 

$$\mu_n = -1.913 imes \mu_N$$
 where,  $\mu_N = rac{e\hbar}{2m_p}$ 

What does the minus mean?

Boothroyd – Sec 1.2 (p 4)





### Definitions: Neutron spin-1/2

spin-<sup>1</sup>/<sub>2</sub> particles

 $S = \frac{1}{2}$  $m_s = -s to + s in integer steps$ - only two choices

Describes an intrinsic angular momentum,  $\vec{s}$ , that the neutron possesses, with



$$\begin{split} |\vec{s}|^2 &= s(s+1)\hbar^2 = \frac{3}{4}\hbar^2\\ s_z &= m_s\hbar \end{split}$$

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### Definitions: Larmor Precession

Magnetic moments in a field experience a **torque** defined as the rate of change of angular momentum

$$\vec{\tau} = \vec{\mu} \times \vec{B}$$
$$\Rightarrow \frac{d\vec{s}}{dt} = \vec{\mu} \times \vec{B}$$
$$\Rightarrow \frac{d\vec{\mu}}{dt} = \gamma \vec{\mu} \times \vec{B}$$

 $d\vec{\mu}$  always in direction perpendicular to both  $\vec{B}$  and  $\vec{\mu}$  therefore,  $\vec{\mu}$  precesses around  $\vec{B}$  maintaining the angle  $\theta$ 

From definition of the cross product







### Definitions: Larmor precession frequency

The angular frequency of the moment vector is defined as the rate of change of the rotation angle  $\phi$  around the field axis

$$\omega = \frac{d\phi}{dt} \Rightarrow d\phi = \omega dt$$

From the figure,  $d\phi\,$  can also be written as the arc-length  $d\vec{\mu}\,$  divided by component of  $\vec{\mu}\,$  in the plane of rotation

$$\begin{split} d\phi &= \frac{|d\vec{\mu}|}{\mu\sin\theta} = \omega dt\\ \text{Rearranging this we get} \quad \frac{|d\vec{\mu}|}{dt} = \omega\mu\sin\theta = \gamma\mu B\sin\theta\\ &\Rightarrow \omega = \gamma B \end{split}$$

This is the Larmor precession frequency,  $\,\omega_L$  and  $\,\gamma$  is known as the gyromagnetic ratio

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Alvarez and Bloch, Physical Review 57, 111 (1940)

### Measuring the neutron magnetic moment: The first ever Neutron Polarization Analysis Experiment



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### **Definitions: Flipping Ratio**

Instead of "polarization" of the beam, experimentalists often talk in terms of the **neutron flipping ratio** 

This can refer to a device – polarizer or analyser – but also to the sample under investigation

$$\Phi = rac{N_{|\uparrow
angle}}{N_{|\downarrow
angle}}$$

where, *N* is the number of neutrons transmitted or scattered.

$$P = \frac{N_{|\uparrow\rangle} - N_{|\downarrow\rangle}}{N_{|\uparrow\rangle} + N_{|\downarrow\rangle}} = \frac{\Delta N}{N} = \frac{\Phi - 1}{\Phi + 1} \quad \text{and} \quad \Phi = \frac{1 + P}{1 - P}$$

This description of a polarized beam is OK for experiments in which a single quantisation axis is defined: **Longitudinal Polarization Analysis** 



### Worked example: Alvarez and Bloch

In the figure on the right, Alvarez and Bloch plot the polarization of their beam as a function of current in their resonance coil – which then gives the resonant frequency. What was the flipping ratio, and polarizing efficiency of the polarizer and analyser ( $P_p$  and  $P_a$ ) assuming they are the same?

The polarization on resonance is,  $P=\Delta I/I=-0.015$ 

Therefore,

$$\Phi = \frac{1 + (-0.015)}{1 - (-0.015)} = \frac{0.985}{1.015} \simeq 0.97$$

The measured polarization is just the product of the individual efficiencies

$$P = P_p \times P_a = P_{p/a}^2$$
$$\Rightarrow P_{p/a} = \sqrt{|P|} \simeq 0.122$$



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FIG. 5. Neutron resonance dip. The magnet current in arbitrary used is plotted against the fractional change  $(\Delta I/I)$  of the intensity of the neutron beam.



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 Image: www.isis.stfc.ac.uk
 Image: mage: ma









# Nuclear scattering

Nuclear Spin-Dependent Scattering

The scattering length depends on the whether the neutron spin is parallel or antiparallel to the nuclear spin

- $b_+$  neutron/nuclear spins add
- $b_-$  neutron/nuclear spins subtract

$$b_{\rm coh} = \bar{b} = \frac{b_+(I+1) + b_-I}{2I+1}$$
$$b_{\rm inc} = \sqrt{\bar{b}^2 - \bar{b}^2} = \frac{b_+ - b_-}{2I+1}\sqrt{(I(I+1))}$$

Boothroyd – Sec 4.1.2 (p 100)

 $b_{\perp}$ 



#### Boothroyd – Sec 4.2 (p 103)

 $\vec{B}(\vec{r}) = \frac{1}{\sqrt{2\pi^3}} \int \vec{B}(\vec{Q}) \exp(i\vec{Q}.\vec{r}) d\vec{Q}$  $\Rightarrow \nabla.\vec{B}(\vec{r}) = \frac{1}{\sqrt{2\pi^3}} \int i\vec{Q}.\vec{B}(\vec{Q}) \exp(i\vec{Q}.\vec{r}) d\vec{Q} = 0$  $\Rightarrow \vec{Q}.\vec{B}(\vec{Q}) = 0$ 

### Reminder: Magnetic scattering

#### **Magnetic scattering**

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The magnetic scattering is a vector interaction –  $V_m = -\vec{\mu}_n . \vec{B}(r)$ 

Since the magnetic field in the sample is divergence-free ( $\nabla . \vec{B}(r) = 0$ ) after taking the Fourier transform we get,

$$V_{m}(Q) = -\vec{\mu_{n}} \cdot \vec{B_{\perp}}(Q)$$
  
=  $-\mu_{0}\vec{\mu_{n}} \cdot \vec{M_{\perp}}(Q)$   
where,  $\vec{M_{\perp}}(Q) = \vec{M}(Q) - (\vec{M}(Q) \cdot \hat{Q})\hat{Q}$   
 $\vec{M_{\perp}}(Q)$   
 $\vec{M_{\perp}}(Q)$ 

#### Neutron probes component of the magnetization perpendicular to Q

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### Magnetic Spin-dependent Scattering (classical)

Magnetic scattering has complex spin-dependence

 $\vec{\tau} = \vec{\mu} \times \vec{B}$ 

There will be a classical torque on the neutron moment if there are components of the field perpendicular to it

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Non-Spin-Flip Spin-flip  $U^{\downarrow\uparrow}$  $U^{\uparrow\downarrow}$  $U^{\downarrow\downarrow}$ • if  $\vec{\mu}$  parallel to  $\vec{B}$  $\vec{Q}$ No torque – so all scattering is non-spin-flip • if  $\vec{\mu}$  perpendicular to  $\vec{B}$ Neutron spin can precess in the field - spin-flip scattering @isisneutronmuon Science and

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Moon, Riste and Koehler, Physical Review 181, 920 (1969)

### Polarized neutron scattering

The polarization-dependent scattering cross-sections can be calculated from the following **Moon-Riste-Koehler** equations

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$$\begin{pmatrix} \frac{d\sigma}{d\Omega} \end{pmatrix}_{ss'} = |\sum_{n} U_{n}^{ss'} \exp\left(i\vec{Q}.\vec{r}_{n}\right)|^{2}$$

$$U^{\uparrow\uparrow} = \bar{b} - pM_{\perp z} + BI_{z} \\ U^{\downarrow\downarrow} = \bar{b} + pM_{\perp z} - BI_{z} \end{pmatrix} U^{\text{NSF}}$$

$$\overline{b} = \frac{(I+1)b_{+} + B}{2I+1} \\ B = \frac{b_{+} - b_{-}}{2I+1} \\ B = \frac{b_{+} - b_{-}}{2I+1} \\ U^{\uparrow\downarrow} = -p\left(M_{\perp x} + iM_{\perp y}\right) + B\left(I_{x} + iI_{y}\right) \\ U^{\downarrow\uparrow} = -p\left(M_{\perp x} - iM_{\perp y}\right) + B\left(I_{x} - iI_{y}\right) \end{pmatrix} U^{\text{SF}}$$

$$p = 1.913 \frac{\mu_{0}}{4\pi} \frac{e^{2}}{m_{e}}$$

where,

- The z-direction is along the direction of the neutron polarization
- *I* is the nuclear spin quantum number and *I<sub>x</sub>*, *I<sub>y</sub>* and *I<sub>z</sub>* are the components of the nuclear spin angular momentum
- $M_{\perp x}$  ,  $M_{\perp y}$  , and  $M_{\perp z}$  are the components of the magnetic field perpendicular to the scattering vector



Boothroyd - Sec 4.4 (p 110)

 $b_{-}$ 

### Coherent (nuclear) Scattering

Assuming no magnetism in the sample – we can work out the **coherent** (see A Boothroyd lectures) nuclear scattering, just by taking the average of the amplitudes

$$\overline{U^{\rm NSF}} = \overline{\overline{b}} \pm \overline{BI_z}$$
$$\overline{U^{\rm SF}} = \overline{BI_x \pm BiI_y}$$

Now, if the nuclear spins are randomly oriented, then the average over any one component of I is zero. Therefore

$$\overline{U^{\rm NSF}} = \overline{b}$$
$$\overline{U^{\rm SF}} = 0$$

#### All the nuclear coherent scattering appears in the Non-Spin-Flip measurements





#### Boothroyd – Sec 4.5.1 (p 112)

### Incoherent (nuclear) Scattering

Again, assuming no magnetism we can work out the **incoherent** scattering which (from A Boothroyd's lectures), is given by the variance in the scattering amplitudes

$$\overline{U^2} - \overline{U}^2$$

Applying this to the NSF amplitudes, we get

$$\overline{(\overline{b} + BI_z)^2} - \overline{(\overline{b} + BI_z)}^2$$
$$= \overline{b^2} + \overline{B^2 I_z^2} + 2\overline{bBI_z} - \overline{(b + BI_z)}^2$$

 $B^{2}I_{z}^{2}$ 

Again, we assume that the nuclei are randomly oriented

Nuclear isotope incoherent

Nuclear spin incoherent

Note that  $\mathbf{I}^2 = I(I+1) = I_x^2 + I_y^2 + I_z^2$ , so for isotropic nuclear spins  $I_x^2 = I_y^2 = I_z^2 = 1/3 I(I+1)$ 

$$\left(\overline{U^2} - \overline{U}^2\right)^{\text{NSF}} = \overline{b^2} - \overline{b}^2 + \frac{1}{3}B^2I(I+1)$$

 $b^{\overline{2}}-\overline{b}^{2}$ 

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Boothroyd – Sec 4.5.1 (p 112)

### Incoherent (nuclear) Scattering

Again, assuming no magnetism we can work out the **incoherent** scattering which (from A Boothroyd's lectures), is given by the variance in the scattering amplitudes

$$\overline{U^2} - \overline{U}^2$$

Applying this to the SF amplitudes, we get

$$\overline{B^2(I_x + iI_y)^2} - \left(\overline{B(I_x + iI_y)}\right)^2$$
$$= \overline{B^2(I_x^2 + I_y^2)}$$

So assuming randomly oriented nuclear spins again, we get

$$\left(\overline{U^2} - \overline{U}^2\right)^{\rm SF} = \frac{2}{3}B^2I(I+1)$$

which is exactly twice the non-spin-flip spin-incoherent scattering





Boothroyd – Sec 4.5.1 (p 112)

### Non-magnetic polarization analysis

$$\left(\frac{d\sigma}{d\Omega}\right)^{\rm NSF} = \left(\frac{d\sigma}{d\Omega}\right)_{\rm coh} + \left(\frac{d\sigma}{d\Omega}\right)_{\rm II} + \frac{1}{3}\left(\frac{d\sigma}{d\Omega}\right)_{\rm SI}$$
$$\left(\frac{d\sigma}{d\Omega}\right)^{\rm SF} = \frac{2}{3}\left(\frac{d\sigma}{d\Omega}\right)_{\rm SI}$$

- if the scattering is fully coherent/isotope incoherent then the scattered polarization is preserved
- if the scattering is fully spin-incoherent then,

$$\vec{P}' = -1/3\vec{P}$$





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Arbe et al., Phys Rev Res 2, 022015 (2020)

### Example 1 – Water ( $D_2O$ ) at 295 K

Measured on LET at ISIS Single-particle (incoherent) and structural (coherent) dynamics are **separated** 





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Nidriche et al., PRX Life **2**, 013005 (2024)

Example 2 – deuterated Green fluorescent protein



New experiments challenging assumptions about "dominant" scattering on deuteration



Boothroyd – Sec 7.3.2 (p 233)

### Scattering from Ferromagnets

Assuming no spin-incoherent scattering in a ferromagnetic material, and assuming that the magnetization is along the polarization direction – we only need to look at the  $U^{\uparrow\uparrow}$  and  $U^{\downarrow\downarrow}$  amplitudes

 $egin{aligned} |U^{\uparrow\uparrow}|^2 &= b^2(Q) + p^2 M_{\perp}^2(Q) - 2b(Q) p M_{\perp}(Q) \ |U^{\downarrow\downarrow}|^2 &= b^2(Q) + p^2 M_{\perp}^2(Q) + 2b(Q) p M_{\perp}(Q) \end{aligned}$ 

Writing this in terms of cross-sections and structure factors

 $\left(\frac{d\sigma}{d\Omega}\right)^{\uparrow\uparrow} \propto |F_N - F_M|^2$  Leading to a flipping ratio of  $\left(\frac{d\sigma}{d\Omega}\right)^{\downarrow\downarrow} \propto |F_N + F_M|^2$   $\Phi = \left(\frac{1-\phi}{1+\phi}\right)^2$  where,  $\phi = \frac{F_M}{F_N}$ 

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### Example 3 – Heusler alloy, Cu<sub>2</sub>MnAl



• Cubic ferromagnet

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• Used as a neutron polarizer

If the (111) Bragg peak from  $Cu_2$ MnAl produces a polarization, P = -0.99, what is the ratio of the magnetic and nuclear (111) structure factors?

We can write the ratio of the structure factors,  $\phi$ , as a function of the flipping ratio

$$\Phi = \frac{1+P}{1-P} = \left(\frac{1-\phi}{1+\phi}\right)^2$$
$$\Rightarrow \phi = \frac{1-\sqrt{\Phi}}{1+\sqrt{\Phi}}$$
The flipping ratio is, 
$$\Phi = \frac{1+(-0.99)}{1-(-0.99)} \simeq 0.005$$
Therefore 
$$\phi = \frac{1-0.071}{1+0.071} \simeq 0.87$$

So  $F_N$  and  $F_M$  don't have to be that close to get a good polarization





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### Example 4 – nickel

In nickel, the ratio of the magnetic and nuclear structure factors for the (400) Bragg peak is  $\phi = 6 \times 10^{-4}$ . What is the flipping ratio? Compare that with the extra magnetic intensity which would be seen in an unpolarized experiment

$$\phi = 6 \times 10^{-4} \Rightarrow \Phi = \left(\frac{1 - 6 \times 10^{-4}}{1 + 6 \times 10^{-4}}\right)^2 = 0.997$$

The ratio of the magnetic and nuclear intensities in an unpolarized experiment is

$$\frac{F_M^2}{F_N^2} = \phi^2 = 3.6 \times 10^{-7} \Rightarrow \frac{\Delta I}{I} = \frac{F_N^2 + F_M^2}{F_N^2} = 1 + \phi^2 = 1.00000036$$

So – magnetic structure factors much easier to measure using polarized neutrons (also, much less sensitive to systematic errors)





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### Scattering from disordered magnets

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Paramagnetic scattering

$$\left(\frac{d\sigma}{d\Omega}\right)^{\rm NSF} = \frac{1}{2} \left[1 - \left(\hat{P} \cdot \hat{Q}\right)^2\right] \left(\frac{d\sigma}{d\Omega}\right)_{\rm mag}$$
$$\left(\frac{d\sigma}{d\Omega}\right)^{\rm SF} = \frac{1}{2} \left[1 + \left(\hat{P} \cdot \hat{Q}\right)^2\right] \left(\frac{d\sigma}{d\Omega}\right)_{\rm mag}$$

We can therefore write the polarization as *H* 

as 
$$P = \frac{\left[1 - \left(\hat{P} \cdot \hat{Q}\right)^2\right] - \left[1 + \left(\hat{P} \cdot \hat{Q}\right)^2\right]}{\left[1 - \left(\hat{P} \cdot \hat{Q}\right)^2\right] + \left[1 + \left(\hat{P} \cdot \hat{Q}\right)^2\right]}$$

This is easily simplified to give **the Halpern-Johnson Equation**  $\vec{P}' = -(\vec{P} \cdot \hat{Q})\hat{Q}$ 





### "Planar multidetector" – XYZ Polarization Analysis

Paramagnetic scattering

$$\left(\frac{d\sigma}{d\Omega}\right)^{\rm NSF} = \frac{1}{2} \left[1 - \left(\hat{P} \cdot \hat{Q}\right)^2\right] \left(\frac{d\sigma}{d\Omega}\right)_{\rm mag} \qquad \left(\frac{d\sigma}{d\Omega}\right)^{\rm SF} = \frac{1}{2} \left[1 + \left(\hat{P} \cdot \hat{Q}\right)^2\right] \left(\frac{d\sigma}{d\Omega}\right)_{\rm mag}$$





### XYZ polarization analysis – paramagnetic scattering

Polarization along x-direction:  $\left(\hat{P}_x \cdot \hat{Q}\right)^2 = \begin{bmatrix} \begin{pmatrix} 1\\0\\0 \end{pmatrix} \cdot \begin{pmatrix} \cos \alpha\\\sin \alpha\\0 \end{bmatrix} \end{bmatrix}^2 = \cos^2 \alpha \quad \begin{cases} \Rightarrow \left(\frac{d\sigma}{d\Omega}\right)_x^{\text{NSF}} = \frac{1 - \cos^2 \alpha}{2} \left(\frac{d\sigma}{d\Omega}\right)_{\text{mag}} = \frac{\sin^2 \alpha}{2} \left(\frac{d\sigma}{d\Omega}\right)_{\text{mag}} \\ \Rightarrow \left(\frac{d\sigma}{d\Omega}\right)^{\text{SF}} = \frac{1 + \cos^2 \alpha}{2} \left(\frac{d\sigma}{d\Omega}\right)_{\text{mag}} \end{cases}$  $\begin{array}{l} \text{arization along y-direction:} \\ \left(\hat{P}_{y}\cdot\hat{Q}\right)^{2} = \left[\begin{pmatrix}0\\1\\0\end{pmatrix}\cdot\begin{pmatrix}\cos\alpha\\\sin\alpha\\0\end{pmatrix}\right]^{2} = \sin^{2}\alpha \quad \begin{cases} \Rightarrow \left(\frac{d\sigma}{d\Omega}\right)_{y}^{\text{NSF}} = \frac{1-\sin^{2}\alpha}{2}\left(\frac{d\sigma}{d\Omega}\right)_{\text{mag}} = \frac{\cos^{2}\alpha}{2}\left(\frac{d\sigma}{d\Omega}\right)_{\text{mag}} \\ \Rightarrow \left(\frac{d\sigma}{d\Omega}\right)^{\text{SF}} = \frac{1+\sin^{2}\alpha}{2}\left(\frac{d\sigma}{d\Omega}\right)_{\text{mag}} \end{cases}$ Polarization along y-direction: arization along z-direction:  $\left(\hat{P}_{y} \cdot \hat{Q}\right)^{2} = \begin{bmatrix} \begin{pmatrix} 0\\0\\1 \end{pmatrix} \cdot \begin{pmatrix} \cos \alpha\\\sin \alpha\\0 \end{bmatrix} \end{bmatrix}^{2} = 0 \qquad \begin{cases} \Rightarrow \left(\frac{d\sigma}{d\Omega}\right)_{z}^{\text{NSF}} = \frac{1}{2} \left(\frac{d\sigma}{d\Omega}\right)_{\text{mag}} \\ \Rightarrow \left(\frac{d\sigma}{d\Omega}\right)^{\text{SF}} = \frac{1}{2} \left(\frac{d\sigma}{d\Omega}\right)_{\text{mag}} \end{cases}$ Polarization along z-direction:

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Schärpf and Capellmann, Phys. Stat. Sol. (a) 135 359 (1993)

#### Boothroyd – Sec 4.6 (p 118)

### XYZ Polarization analysis – Schärpf Equations

$$\begin{split} \left(\frac{d\sigma}{d\Omega}\right)_{x}^{\text{NSF}} &= \left(\frac{d\sigma}{d\Omega}\right)_{\text{coh+II}} + \frac{1}{3}\left(\frac{d\sigma}{d\Omega}\right)_{\text{SI}} + \frac{\sin^{2}\alpha}{2}\left(\frac{d\sigma}{d\Omega}\right)_{\text{mag}} \\ \left(\frac{d\sigma}{d\Omega}\right)_{y}^{\text{NSF}} &= \left(\frac{d\sigma}{d\Omega}\right)_{\text{coh+II}} + \frac{1}{3}\left(\frac{d\sigma}{d\Omega}\right)_{\text{SI}} + \frac{\cos^{2}\alpha}{2}\left(\frac{d\sigma}{d\Omega}\right)_{\text{mag}} \\ \left(\frac{d\sigma}{d\Omega}\right)_{z}^{\text{NSF}} &= \left(\frac{d\sigma}{d\Omega}\right)_{\text{coh+II}} + \frac{1}{3}\left(\frac{d\sigma}{d\Omega}\right)_{\text{SI}} + \frac{1}{2}\left(\frac{d\sigma}{d\Omega}\right)_{\text{mag}} \\ \left(\frac{d\sigma}{d\Omega}\right)_{x}^{\text{SF}} &= \frac{2}{3}\left(\frac{d\sigma}{d\Omega}\right)_{\text{SI}} + \frac{1 + \cos^{2}\alpha}{2}\left(\frac{d\sigma}{d\Omega}\right)_{\text{mag}} \\ \left(\frac{d\sigma}{d\Omega}\right)_{y}^{\text{SF}} &= \frac{2}{3}\left(\frac{d\sigma}{d\Omega}\right)_{\text{SI}} + \frac{1 + \sin^{2}\alpha}{2}\left(\frac{d\sigma}{d\Omega}\right)_{\text{mag}} \\ \left(\frac{d\sigma}{d\Omega}\right)_{z}^{\text{SF}} &= \frac{2}{3}\left(\frac{d\sigma}{d\Omega}\right)_{\text{SI}} + \frac{1}{2}\left(\frac{d\sigma}{d\Omega}\right)_{\text{mag}} \end{split}$$



Otto Schärpf (1929-2019)

- 6 equations
- 4 unknowns
- $\alpha$  known as the Schärpf Angle



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### XYZ Polarization analysis – Schärpf Equations

Coherent (+II) nuclear scattering

$$\left(\frac{d\sigma}{d\Omega}\right)_{\rm coh+II} = \frac{1}{6} \left[ 2 \left(\frac{d\sigma}{d\Omega}\right)^{\rm TNSF} - \left(\frac{d\sigma}{d\Omega}\right)^{\rm TSF} \right]$$

where TNSF and TSF refer to the total non-spin-flip and spin-flip scattering respectively

Magnetic scattering - can be independently calculated in two ways

$$\left(\frac{d\sigma}{d\Omega}\right)_{\rm mag} = 2\left(\frac{d\sigma}{d\Omega}\right)_{x}^{\rm SF} + 2\left(\frac{d\sigma}{d\Omega}\right)_{y}^{\rm SF} - 4\left(\frac{d\sigma}{d\Omega}\right)_{z}^{\rm SF} \left(\frac{d\sigma}{d\Omega}\right)_{\rm mag} = 4\left(\frac{d\sigma}{d\Omega}\right)_{z}^{\rm NSF} - 2\left(\frac{d\sigma}{d\Omega}\right)_{x}^{\rm NSF} - 2\left(\frac{d\sigma}{d\Omega}\right)_{y}^{\rm NSF}$$

can take the average of these

Spin-incoherent scattering

$$\left(\frac{d\sigma}{d\Omega}\right)_{\rm SI} = \frac{1}{2} \left(\frac{d\sigma}{d\Omega}\right)^{\rm TSF} - \left(\frac{d\sigma}{d\Omega}\right)_{\rm mag}$$





### Stewart, et al., J. Appl. Phys 42, 69 (2009) Why is this necessary?





### Reminder – Rules for Polarized Neutron Scattering

- 1. The nuclear coherent and isotope incoherent scattering is always NSF
- 2. The spin-incoherent scattering is always 1/3 NSF and 2/3 SF
- 3. Magnetism: must be in a direction perpendicular to  $\vec{Q}$  and:
  - NSF if parallel to  $\vec{P}$
  - SF if perpendicular to  $\vec{P}$

 

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### Neutron Spin-Echo

**Precession field** Precession field Polarizer Analyzer  $B_2L_2$  $B_1L_1$  $\pi$  $\pi/2$  $\pi/2$ flippers Detector  $\mathbf{P}$ Number of radians precessed in each arm:  $\phi_L = \omega t = \omega \times \frac{L}{v} = \frac{\gamma BL}{v}$ If  $B_1L_1 = B_2L_2$ , and  $v_1 = v_2$  then the neutron spins will all re-align:  $\Delta \phi_L = \frac{\gamma B_1 L_1}{\gamma B_2 L_2} = 0$ 



Boothroyd - Sec 5.9 (p 176)

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### **Neutron Spin-Echo**



(assumes that the sample doesn't depolarize the beam)

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### NSE Resolution – Spin-Echo Time

Neutron energy is  $E = \frac{1}{2}m_n v^2 \Rightarrow \Delta E = m_n v \Delta v$ 

In NSE we measure a phase shift  $\phi = \frac{\gamma BL}{v} \Rightarrow \Delta \phi = \frac{\gamma BL}{v^2} \Delta v = \frac{\gamma BL \Delta E}{m_n v^3}$ 

Since we know that  $\Delta E = \hbar \omega$  and that the phase shift  $\phi = \omega t$  we can identify the Spin-Echo time

$$t = \frac{\gamma B L \hbar}{m_n v^3} = \frac{\gamma B L m_n^2 \lambda^3}{2\pi h^2}$$

Putting in numbers: BL = 0.27~ Tm on WASP at ILL, and  $\lambda = 1~$  nm we find a Spin-Echo time, t = 18~ns

This is equivalent to  $\omega = 2\pi/t = 0.35\,{
m GHz}$ , or 0.08  $\mu{
m eV}$ 

#### Much higher energy resolution than on a conventional spectrometer



Alles Sum

### Arbe et al., J. Chem. Phys. **158** (2023) 184502 **Example 6 — dynamics in van der Waals liquid**

Tetrahydrofuran – THF, C<sub>4</sub>H<sub>8</sub>O

LET at ISIS

SIS Neutron and

Muon Source



Coherent and incoherent scattering in deuterated dTHF

Fit is a model consisting of diffusion (incoherent) and structural relaxation (coherent) – too complicated to disentangle without polarized neutrons – and which requires NSE for the resolution at low Q

www.isis.stfc.ac.uk

www.isis

WASP at ILL



A les Sun

Mostly incoherent scattering in protonated pTHF



Polarized neutrons can be used to enhance almost any neutron scattering experiment, either by

- 1. Separating and providing additional information on the **sources of scattering** (nuclear, magnetic, incoherent)
- 2. Improving the **resolution**



