

Neutron Magnetic Scattering

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The plan

- How does the neutron interact with magnetism?
- The fundamental rule of neutron magnetic scattering
- Elastic scattering, and how to understand it
- Magnetic form factors
- Generalized susceptibility
- Inelastic scattering
 - Crystal fields and molecular magnets
 - Magnons
 - Spin wave continua
- Critical scattering
- Short-ranged order

How does the neutron interact with magnetism?

Neutrons have no charge, but they do have a magnetic moment.

The magnetic moment is given by the neutron's spin angular momentum:

$$-\gamma\mu_{\text{N}}\hat{\sigma}$$

where:

- γ is a constant (=1.913)
- μ_{N} is the nuclear magneton
- $\hat{\sigma}$ is the quantum mechanical Pauli spin operator

Normally refer to it as a spin-1/2 particle

How does the neutron interact with magnetism?

Through the cross-section!

$$\frac{d^2\sigma}{d\Omega \cdot dE} = \underbrace{\frac{k'}{k} \left(\frac{m_n}{2\pi\hbar^2} \right)^2}_{\text{Conservation of energy}} \underbrace{\sum_{\xi,s} p_{\xi} p_s \sum_{\xi',s'} |\langle \mathbf{k}', s', \xi' | \hat{V}(\mathbf{r}) | \mathbf{k}, s, \xi \rangle|^2}_{\text{Probabilities of initial target state and neutron spin}} \delta(\hbar\omega + E_{\xi} - E_{\xi'})$$

The *matrix element*, which contains all the physics.

$\hat{V}(\mathbf{r})$ is the *pseudopotential*,
which for magnetism is given by:

$$\hat{V}_m(\mathbf{r}) = -\gamma\mu_N \hat{\boldsymbol{\sigma}} \cdot \mathbf{B}(\mathbf{r})$$

where $\mathbf{B}(\mathbf{r})$ is the magnetic induction.

G. L. Squires, *Introduction to the theory of thermal neutron scattering*, Dover Publications, New York, 1978
 W. Marshall and S. W. Lovesey, *Theory of thermal neutron scattering*, Oxford University Press, Oxford, 1971
 S. W. Lovesey, *Theory of neutron scattering from condensed matter*, Oxford University Press, Oxford, 1986

Elastic scattering

If the incident neutron energy = the final neutron energy, the scattering is *elastic*.

$$\frac{d\sigma}{d\Omega} = \left(\frac{m_n}{2\pi\hbar^2} \right)^2 \sum_{s'} p_s \left| \langle \mathbf{k}', s' | \hat{V}(\mathbf{r}) | \mathbf{k}, s \rangle \right|^2$$

Forget about the spins for the moment (*unpolarized* neutron scattering) and integrate over all $\mathbf{r}$:

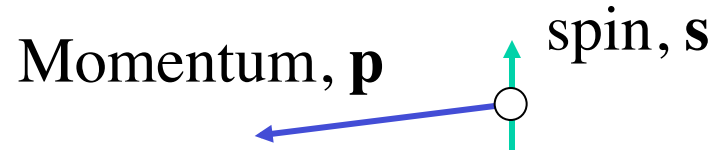
$$\langle \mathbf{k}' | \hat{V}(\mathbf{r}) | \mathbf{k} \rangle = \int \hat{V}(\mathbf{r}) e^{i\mathbf{Q} \cdot \mathbf{r}} \cdot d\mathbf{r}$$

Momentum transfer $\mathbf{Q} = \mathbf{k} - \mathbf{k}'$

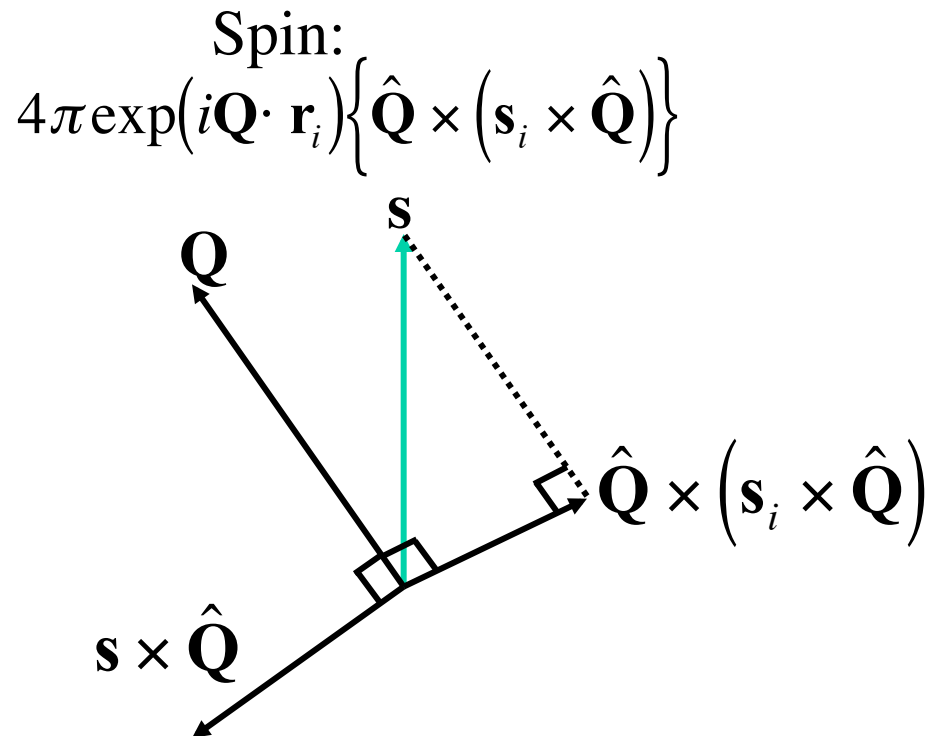
The elastic cross-section is then directly proportional to the *Fourier transform squared* of the potential.
Neutron scattering thus works in Fourier space, otherwise called *reciprocal space*.

Learn about Fourier transforms!

Magnetism is caused by unpaired electrons or movement of charge.

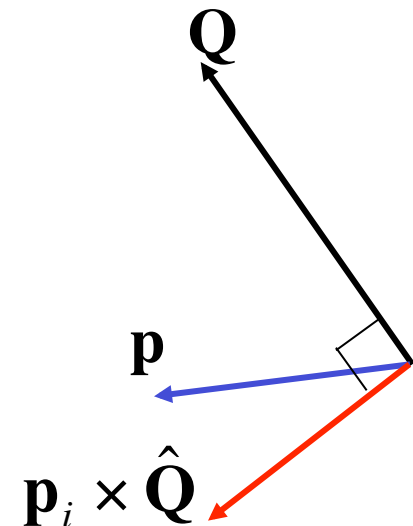


$$\langle \mathbf{k}' | \hat{V}_m(\mathbf{r}_i) | \mathbf{k} \rangle =$$



Movement / Orbital

$$\frac{4\pi i}{\hbar Q} \exp(i\mathbf{Q} \cdot \mathbf{r}_i) (\mathbf{p}_i \times \hat{\mathbf{Q}})$$



Neutrons *only ever* see the
components of the magnetization
that are *perpendicular* to the
scattering vector!

The fundamental rule of neutron magnetic scattering

Taking elastic scattering again:

$$\frac{d\sigma}{d\Omega} = \left(\frac{m_n}{2\pi\hbar^2} \right)^2 \left| \langle \mathbf{k}' | \hat{V}(\mathbf{r}) | \mathbf{k} \rangle \right|^2$$

Neutron scattering therefore probes the components of the sample magnetization that are *perpendicular* to the neutron's momentum transfer, $\mathbf{Q}$.

$$\int V_m(\mathbf{r}) e^{i\mathbf{Q}\cdot\mathbf{r}} d\mathbf{r} = \mathbf{M}_\perp(\mathbf{Q})$$

and

$$\frac{d\sigma_{\text{magnetic}}}{d\Omega} = \langle \mathbf{M}_\perp^*(\mathbf{Q}) \rangle \langle \mathbf{M}_\perp(\mathbf{Q}) \rangle$$

Neutron scattering measures the *correlations* in magnetization,
i.e. the influence a magnetic moment has on its neighbours.

It is capable of doing this over all length scales, limited only by wavelength.

Learn your Fourier transforms!

and

Learn and understand the convolution theorem!

$$f(r) \otimes g(r) = \int f(x)g(r-x) \cdot dx$$

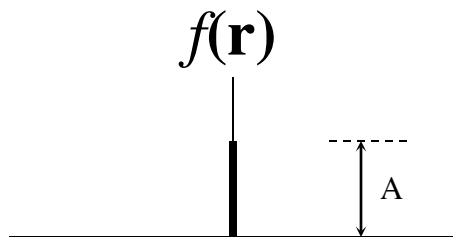
$$\mathfrak{F}(f(r)) = F(q)$$

$$\mathfrak{F}(g(r)) = G(q)$$

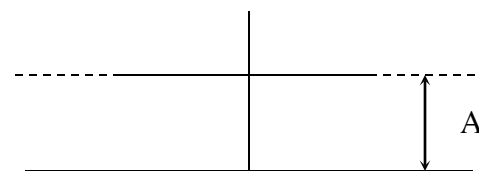
$$\mathfrak{F}(f(r) \otimes g(r)) = F(q) \times G(q)$$

Fourier Transforms $F(\mathbf{Q}) = \int f(\mathbf{r}) e^{i\mathbf{Q} \cdot \mathbf{r}} \cdot d\mathbf{r}$

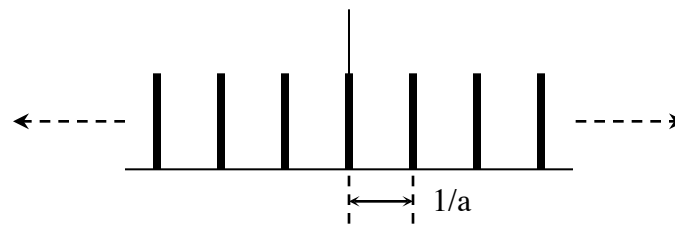
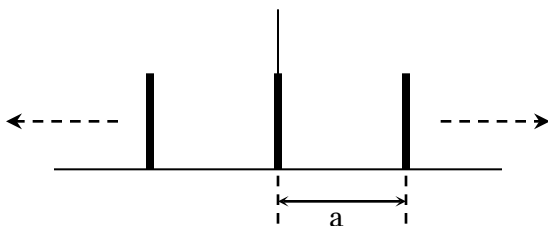
A Delta function



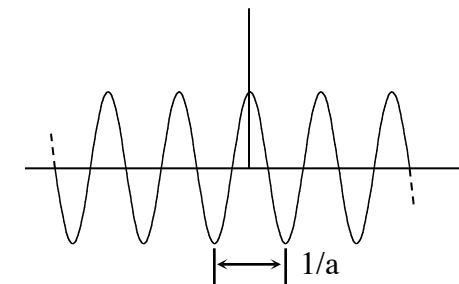
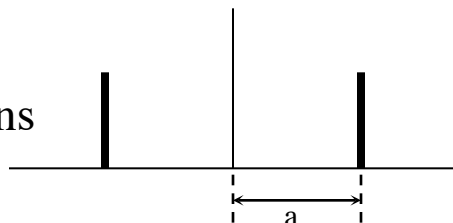
$F(\mathbf{Q})$



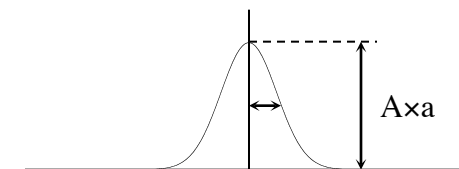
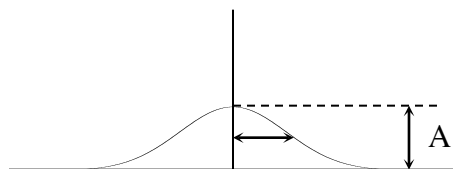
A series of
Delta functions



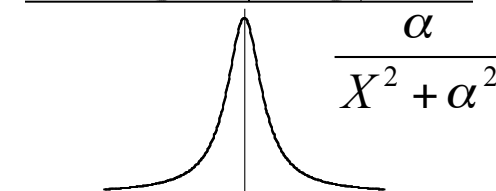
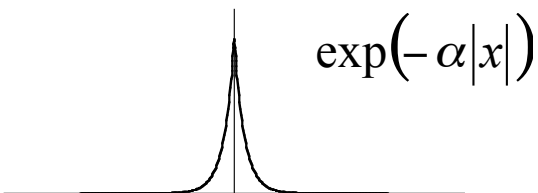
Two Delta functions



A Gaussian



An exponential



Elastic scattering

$$\frac{d\sigma}{d\Omega} = \left(\frac{m_n}{2\pi\hbar^2} \right)^2 \sum_{\xi', s'} p_s \left| \langle \mathbf{k}', s' | \hat{V}(\mathbf{r}) | \mathbf{k}, s \rangle \right|^2$$

$$\propto \underbrace{\int \left| \langle \hat{V} \rangle \right|^2 e^{i\mathbf{Q} \cdot \mathbf{r}} \cdot d\mathbf{r}}_{\text{Long-range order}} + \underbrace{\left(\left| \langle \hat{V}^2 \rangle \right| - \left| \langle \hat{V} \rangle \right|^2 \right) \int z(\mathbf{r}) e^{i\mathbf{Q} \cdot \mathbf{r}} \cdot d\mathbf{r}}_{\text{Short-range order}}$$

The contribution from deviations from
the average structure:
Short-range order

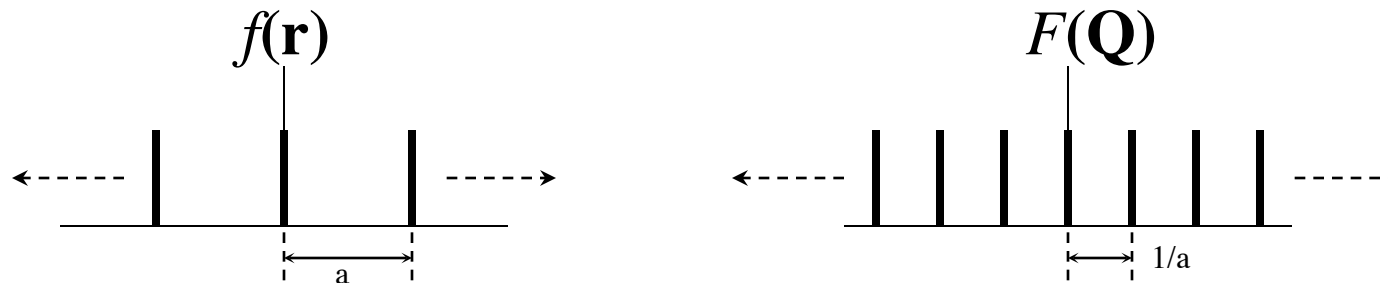
The contribution from the average
structure of the sample:
Long-range order

Magnetic structure determination

Crystalline structures

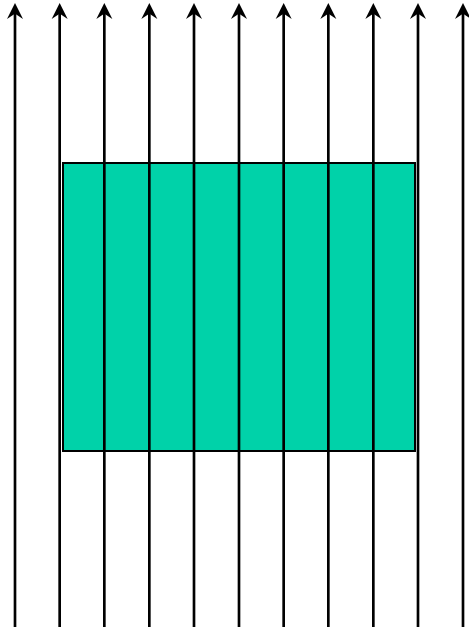
$$\frac{d\sigma}{d\Omega} \propto \int \left| \langle \hat{V} \rangle \right|^2 e^{i\mathbf{Q} \cdot \mathbf{r}} \cdot d\mathbf{r}$$

The Fourier transform from a series of delta-functions

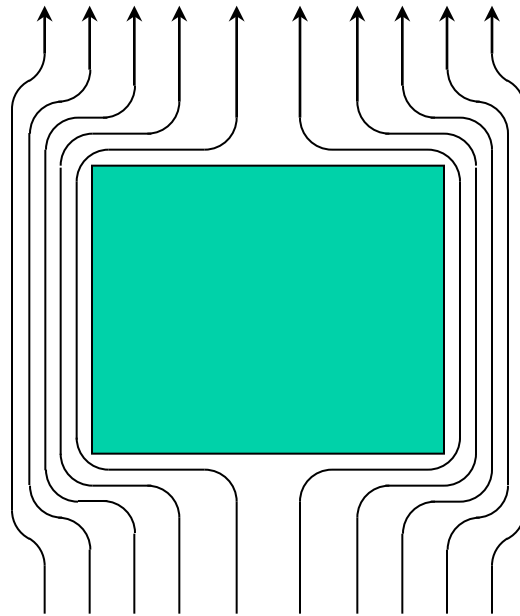


Bragg's Law: $2d\sin\theta = \lambda$
Leads to Magnetic Crystallography

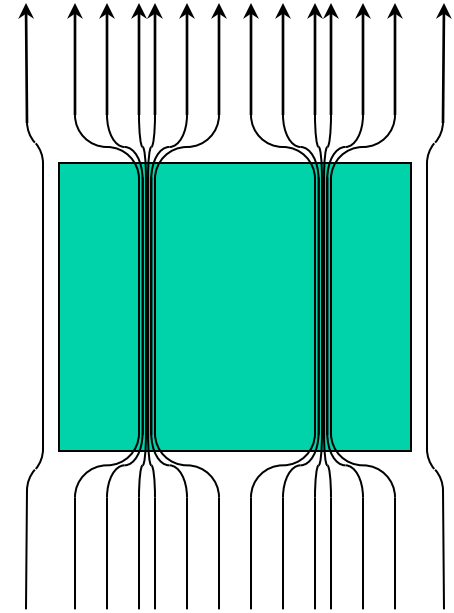
Superconductivity



Normal state



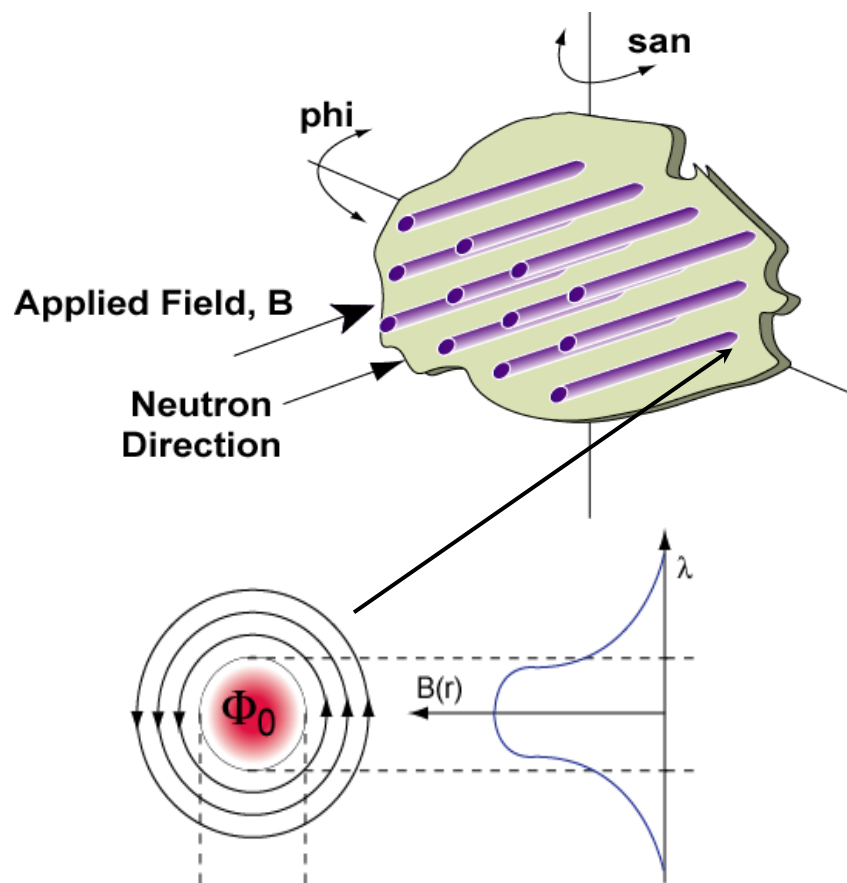
Superconducting state



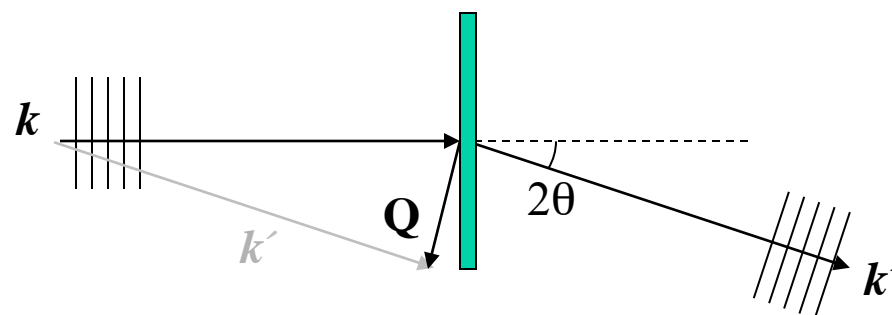
Flux line lattice

A simple example of magnetic elastic scattering

MgB_2 is a superconductor below 39K, and expels all magnetic field lines (Meissner effect).
Above a critical field, flux lines penetrate the sample.



Scattering geometry

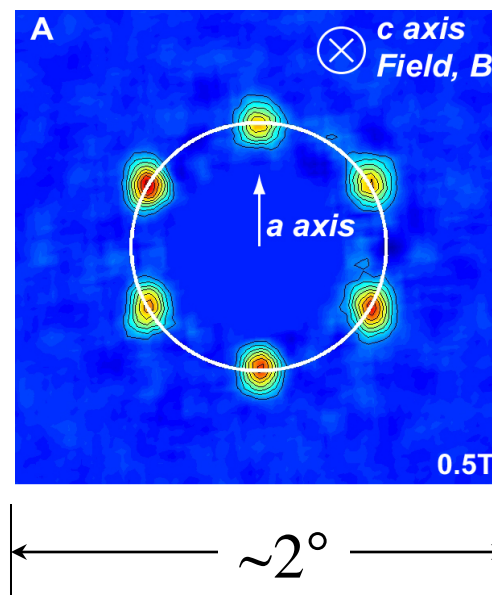
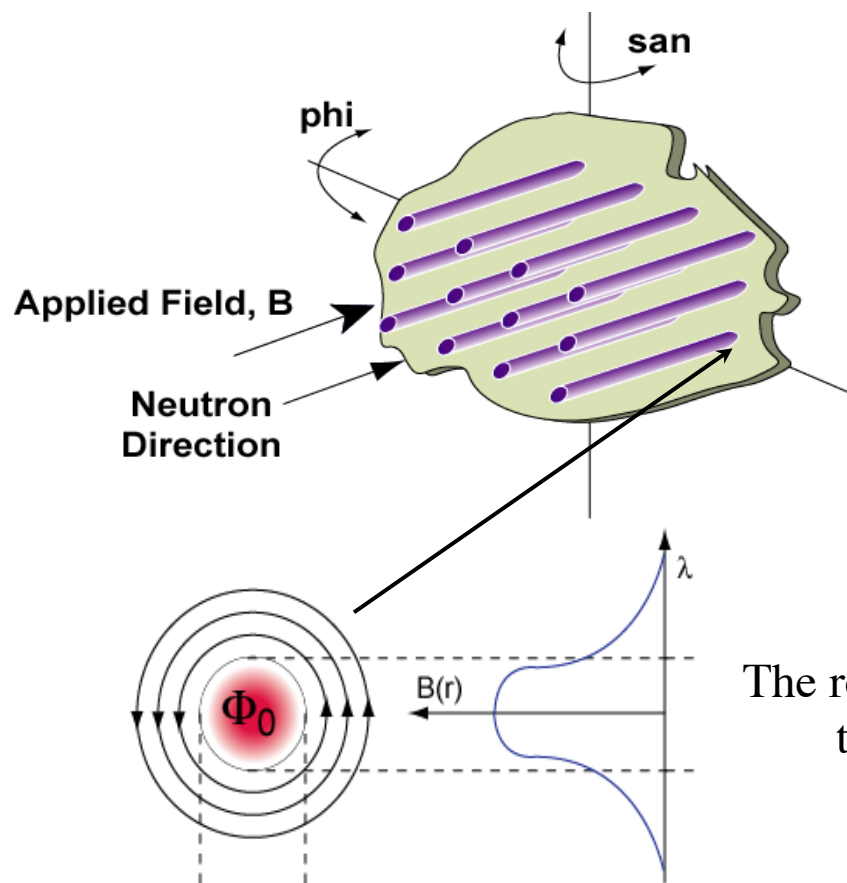


The momentum transfer, $\mathbf{Q}$, is roughly perpendicular to the flux lines, therefore all the magnetization is seen.

$$(\text{recall } \frac{d\sigma_{\text{magnetic}}}{d\Omega} = \langle \mathbf{M}_{\perp}^*(\mathbf{Q}) \rangle \langle \mathbf{M}_{\perp}(\mathbf{Q}) \rangle)$$

A simple example of magnetic elastic scattering

MgB₂ is a superconductor below 39K, and expels all magnetic field lines (Meissner effect).
Above a critical field, flux lines penetrate the sample.



Via Bragg's Law
 $2d\sin\theta = \lambda$
 $\lambda = 10 \text{ \AA}$
 $d = 425 \text{ \AA}$

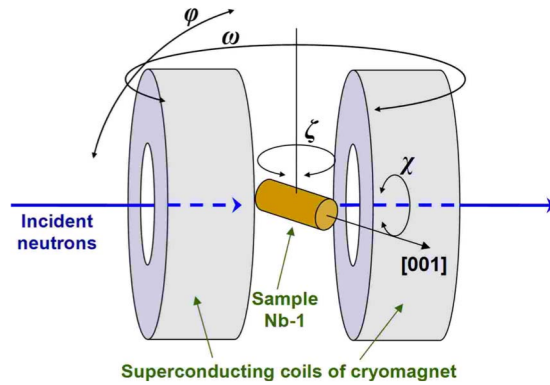
The reciprocal lattice has 60° rotational symmetry,
therefore the flux line lattice is hexagonal

Cubitt et. al. Phys. Rev. Lett. **91** 047002 (2003)

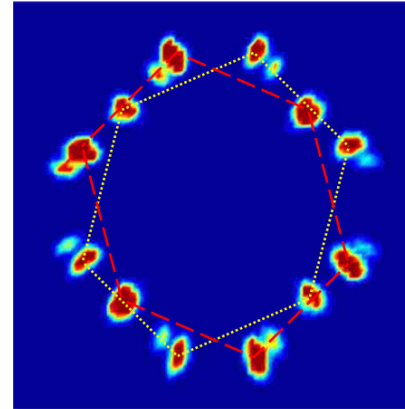
Cubitt et. al. Phys. Rev. Lett. **90** 157002 (2003)

Flux line lattices

Pure Niobium

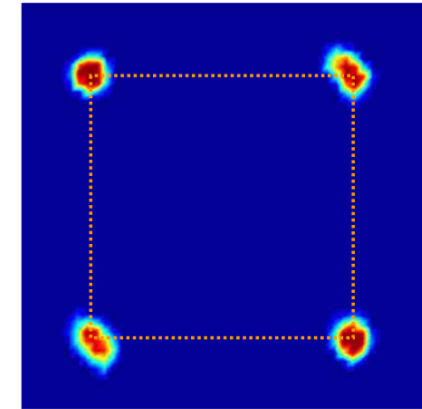


200 mT

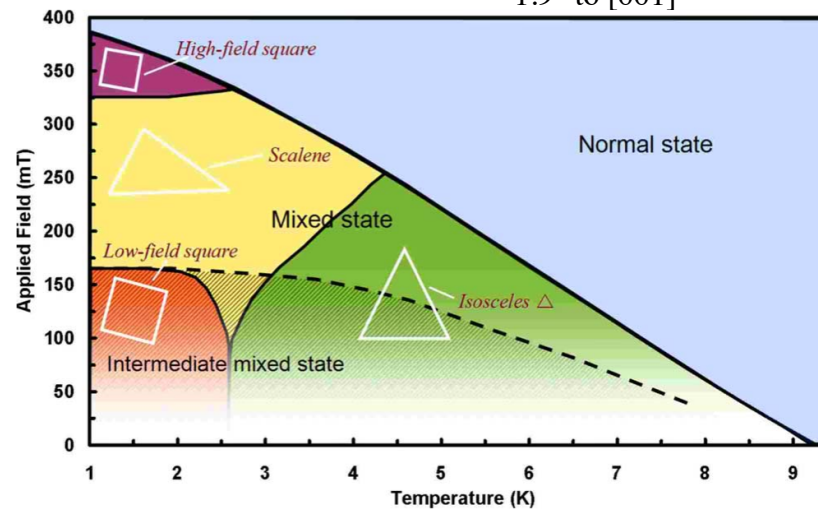


0.1° to (110)
 1.9° to [001]

350 mT



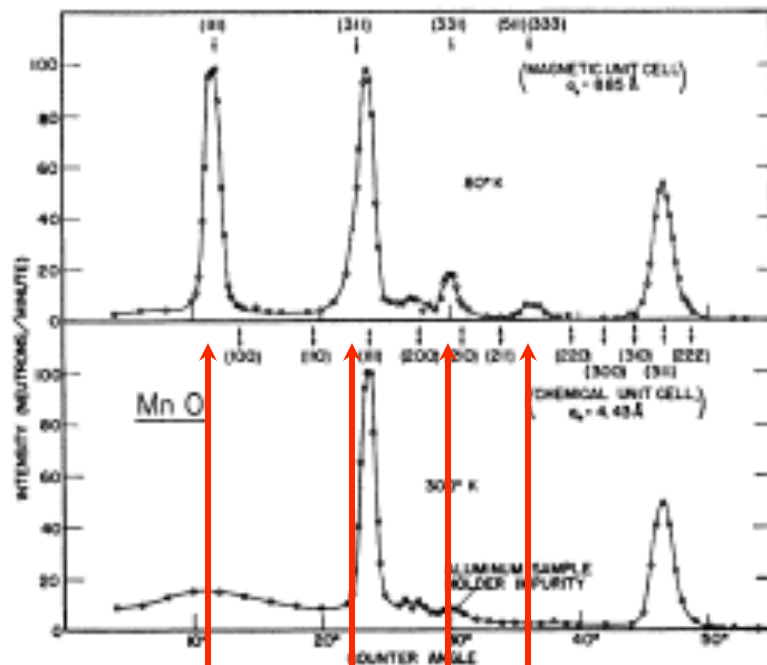
0.1° to (110)
 1.9° to [001]



M. Laver et. al. Phys. Rev. B **79** 014518 (2009)

Antiferromagnetism in MnO

Bragg peaks from crystal structure



80K (antiferromagnetic)

300K (paramagnetic)

New Bragg peaks

C. G. Shull & J. S. Smart, Phys. Rev. **76** (1949) 1256

Antiferromagnetism in MnO

New magnetic Bragg peaks

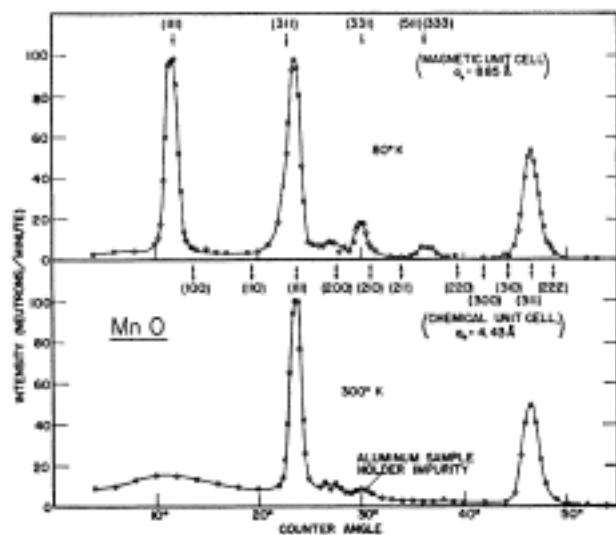
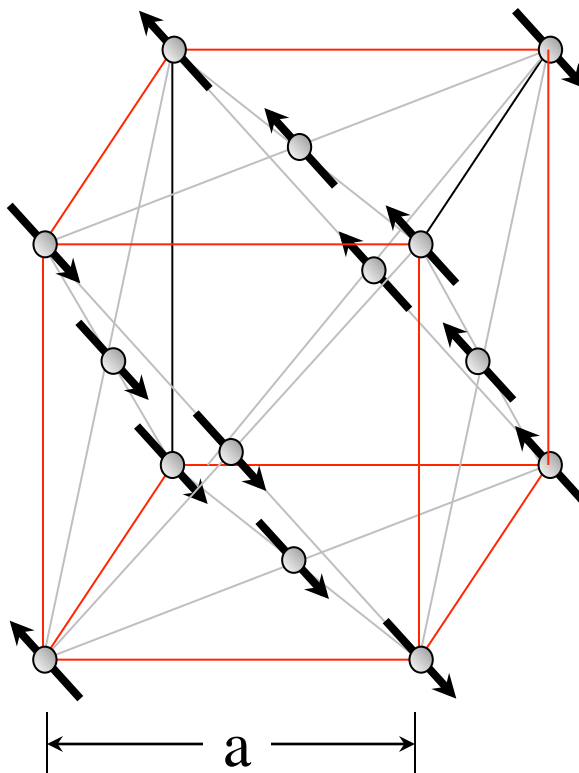


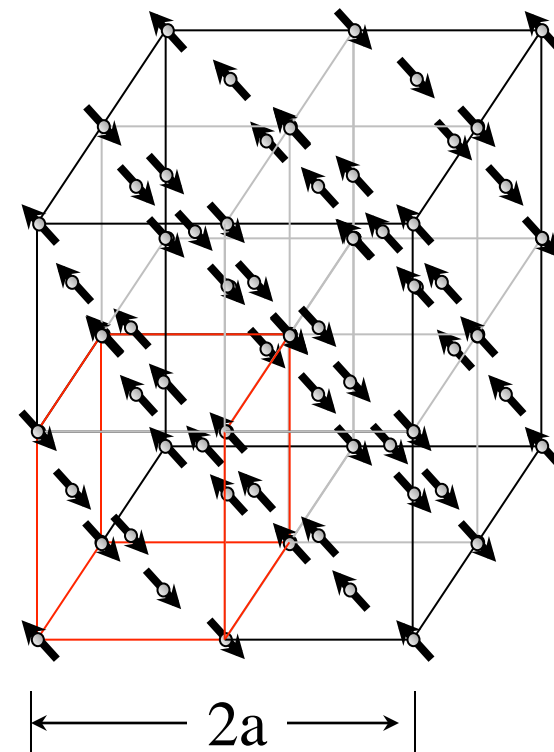
FIG. 1. Neutron diffraction patterns for MnO at room temperature and at 80°K.

C. G. Shull & J. S. Smart, Phys. Rev. **76** (1949) 1256

Magnetic structure



Magnetic unit cell



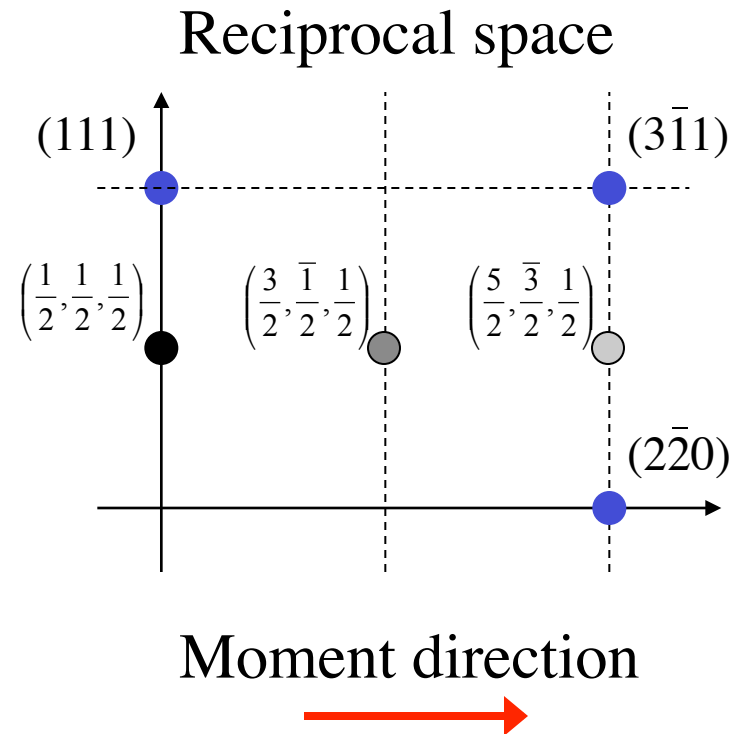
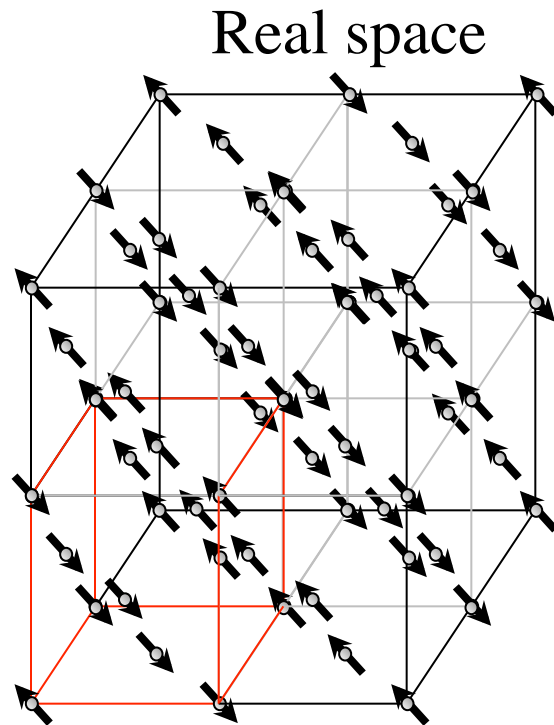
C. G. Shull *et al.*, Phys. Rev. **83** (1951) 333

H. Shaked *et al.*, Phys. Rev. B **38** (1988) 11901

Antiferromagnetism in MnO

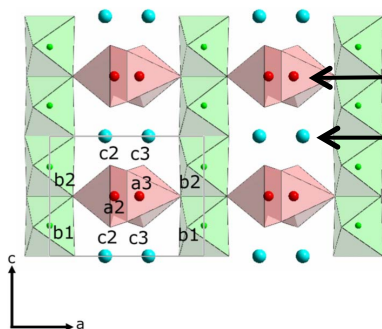
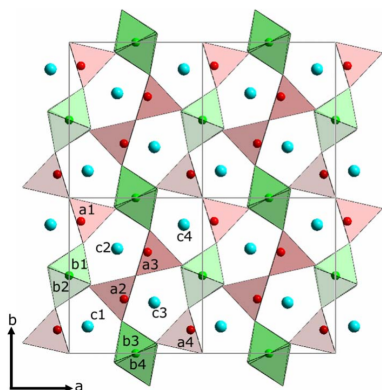
The moments are said to lie in the (111) plane

H. Shaked *et al.*, Phys. Rev. B **38** (1988) 11901



The magnetic structure of YMn_2O_5

Crystal structure

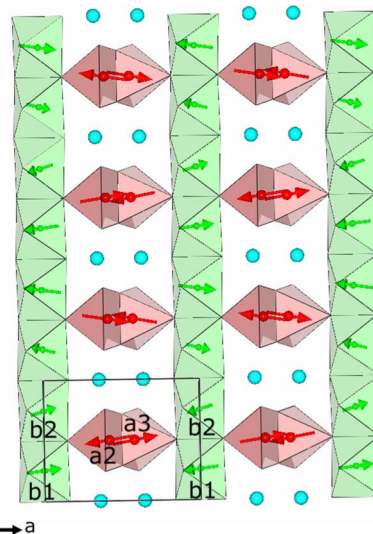
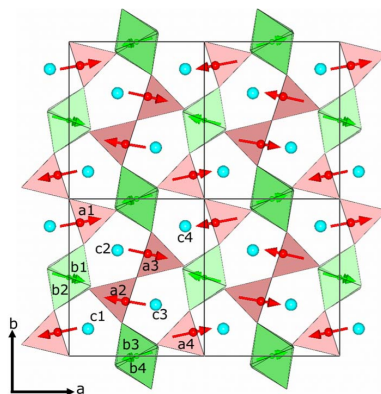


← Mn^{4+}
← Mn^{3+}
← Y / rare earth

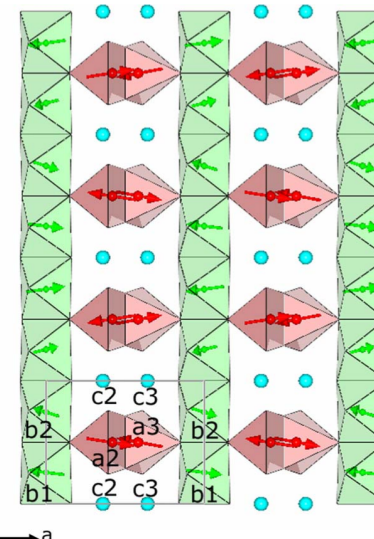
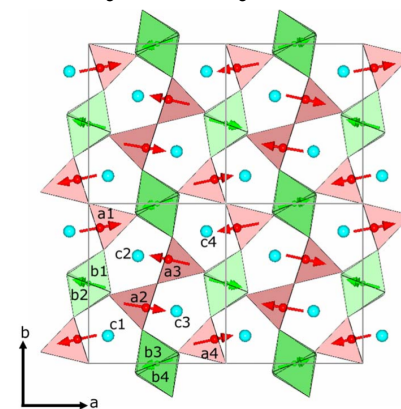
C. Vecchini *et al.*, PRB **77** (2008) 134434

Domain 1

(domains are related by inversion symmetry)



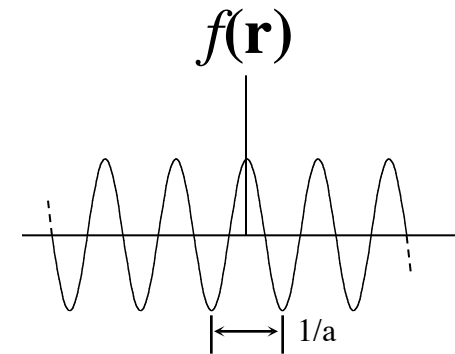
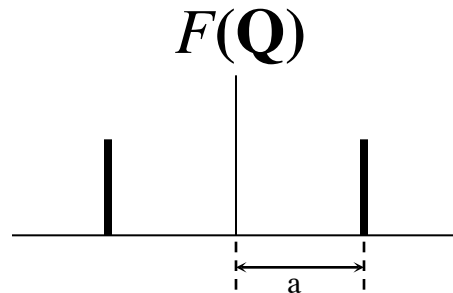
Domain 2



YMn_2O_5 has a *commensurate* magnetic structure

Antiferromagnetism in Chromium

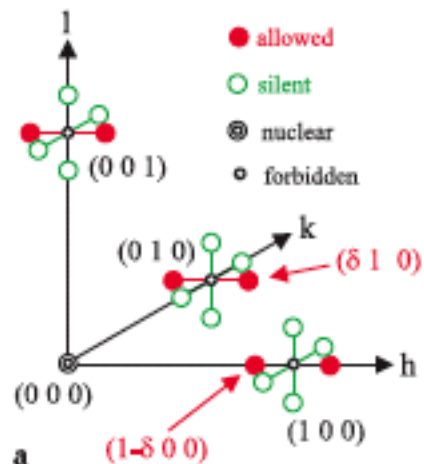
The Fourier Transform for two Delta functions:



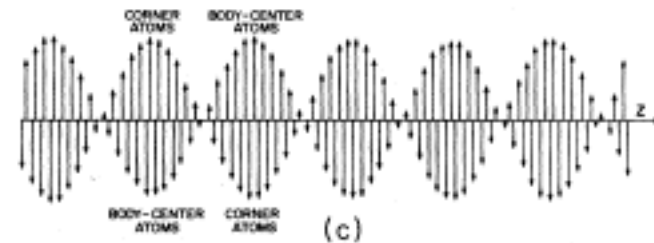
E.Fawcett, Rev. Mod. Phys. **60** (1988) 209

Chromium is an example of an *itinerant* antiferromagnet

Reciprocal space



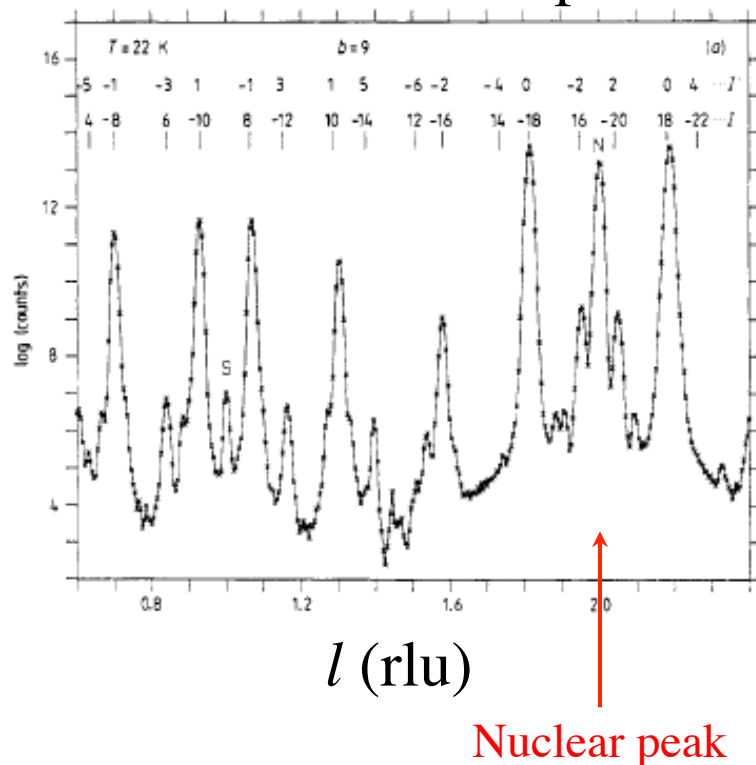
Real space



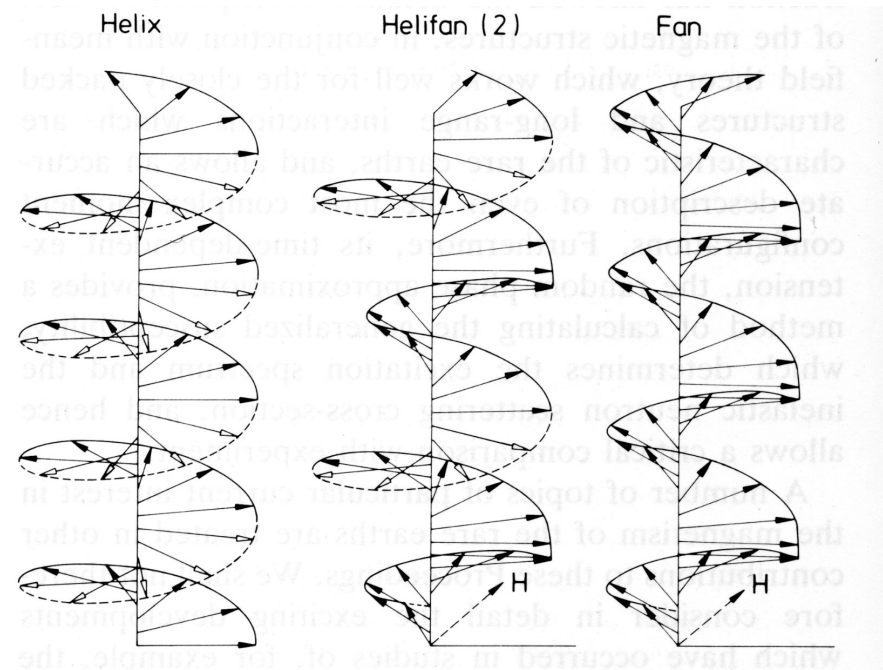
a Spin Density wave

Magnetic structure of Holmium

Scan along $[00l]$,
incommensurate peaks



Real space structures

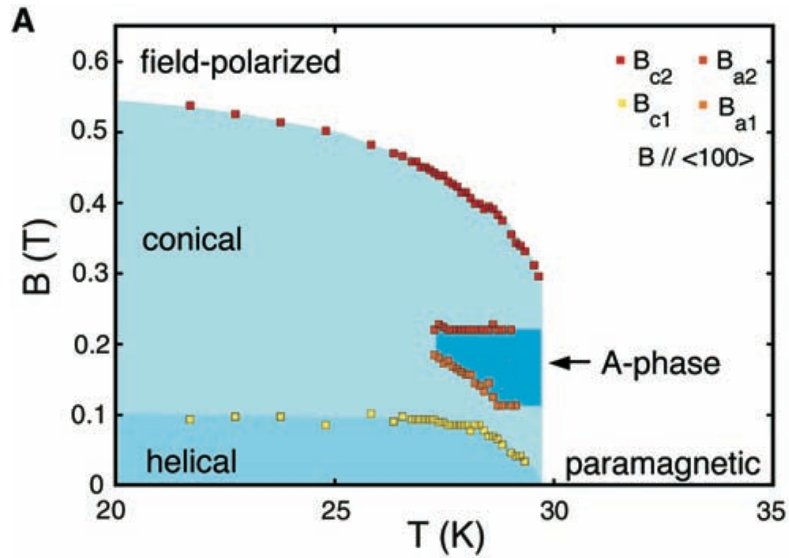


W. C. Koehler, in *Magnetic Properties of Rare Earth Metals*, ed. R. J. Elliot (Plenum Press, London, 1972) p. 81

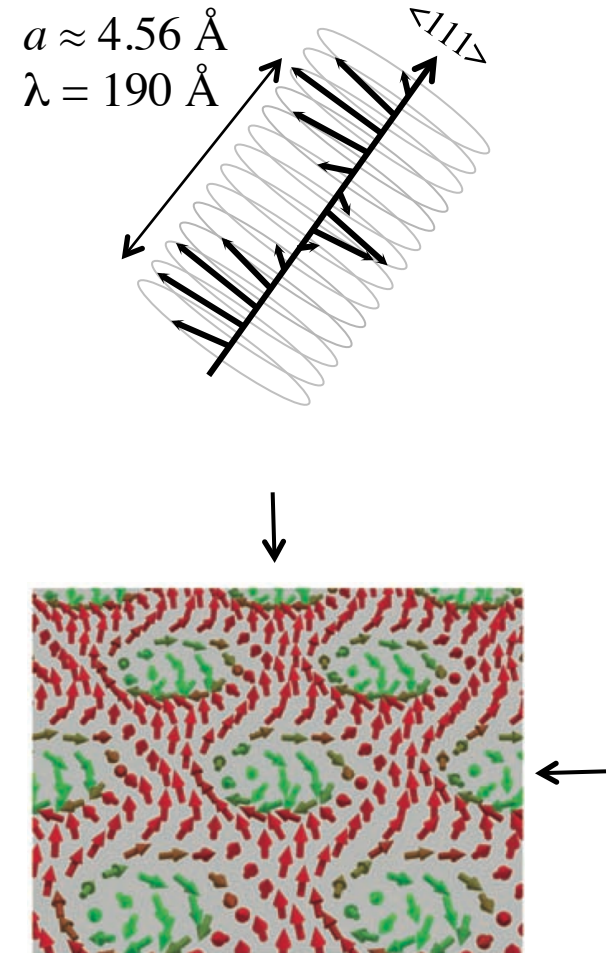
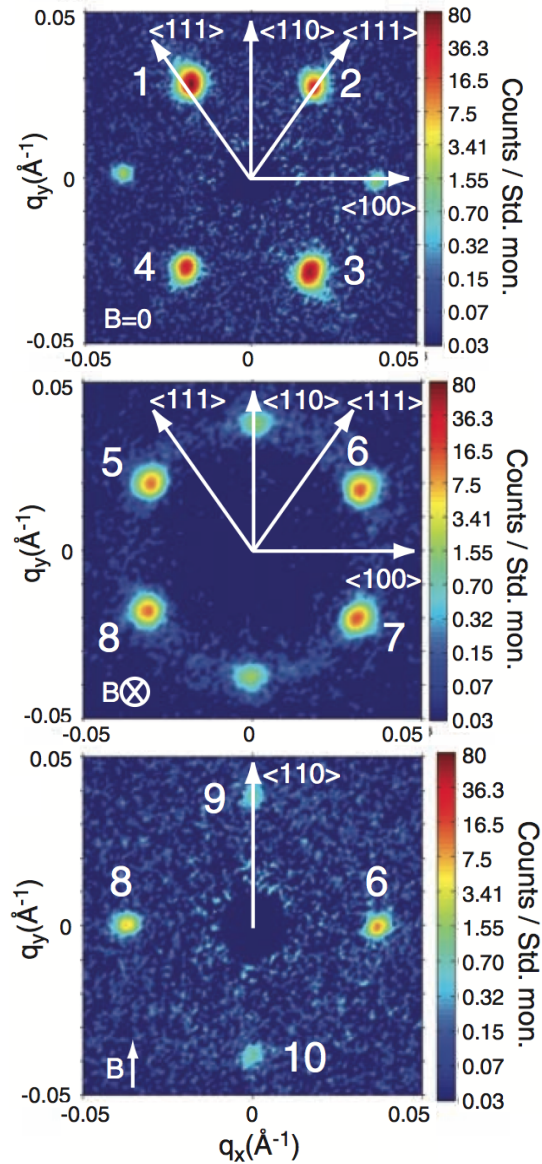
R. A. Cowley and S. Bates, *J. Phys. C* **21** (1988) 4113

A. R. Mackintosh and J. Jensen, *Physica B* **180 & 181** (1992) 1

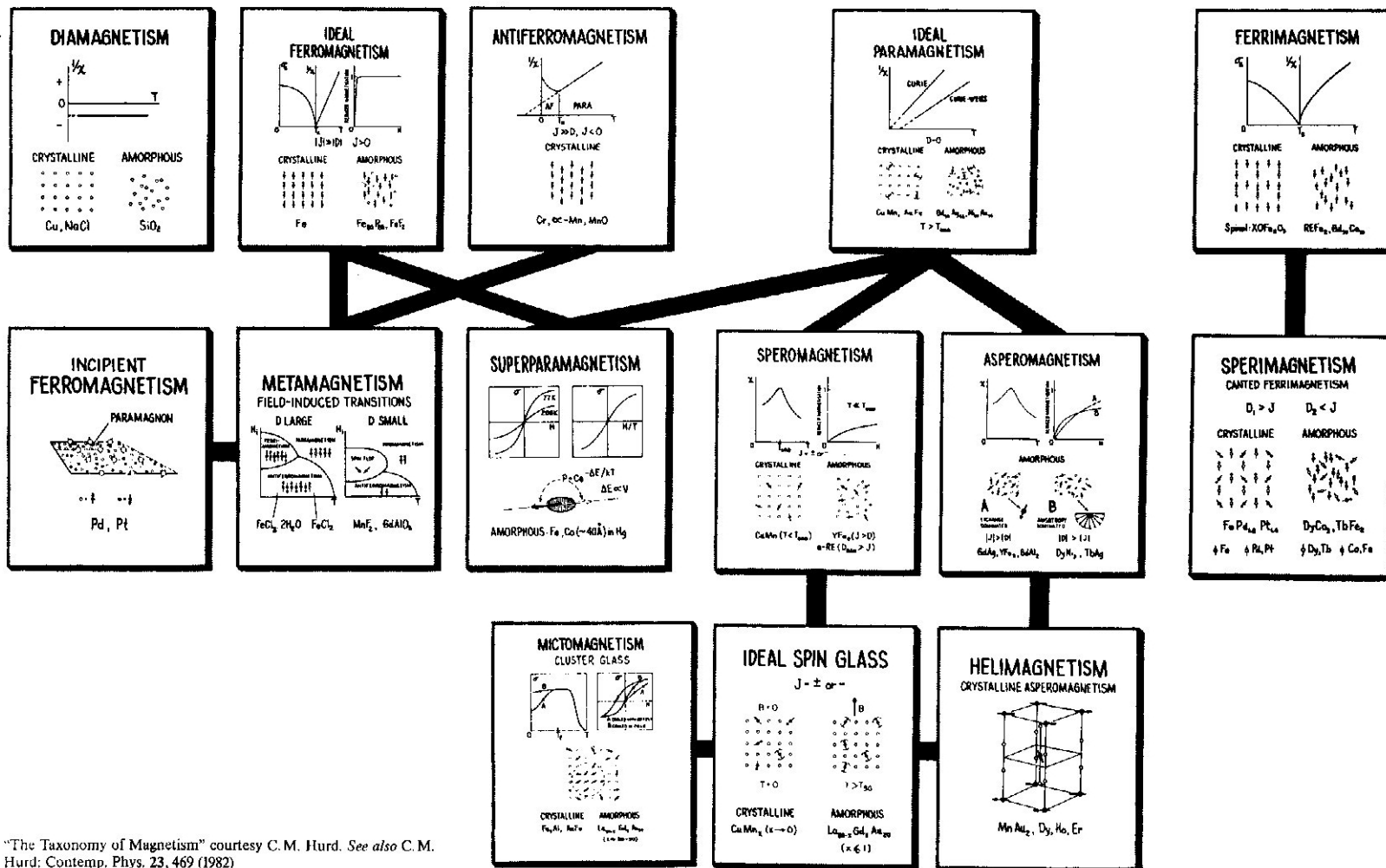
Skyrmions in MnSi



S. Mühlbauer et. al. Science **323** 915 (2009)



The 'Family Tree' of Magnetism

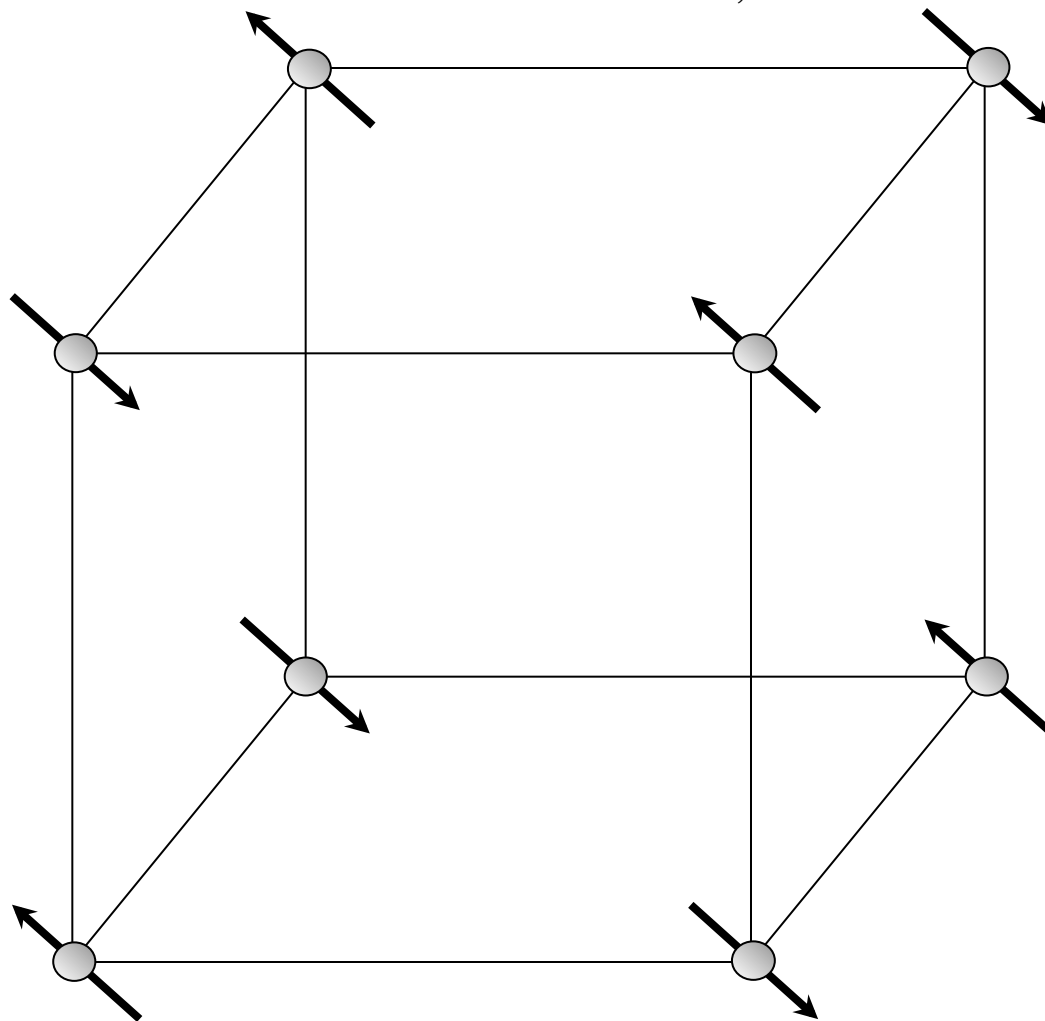


"The Taxonomy of Magnetism" courtesy C.M. Hurd. See also C.M. Hurd; Contemp. Phys. 23, 469 (1982)

Magnetic form factors

A magnetic moment is spread out in space

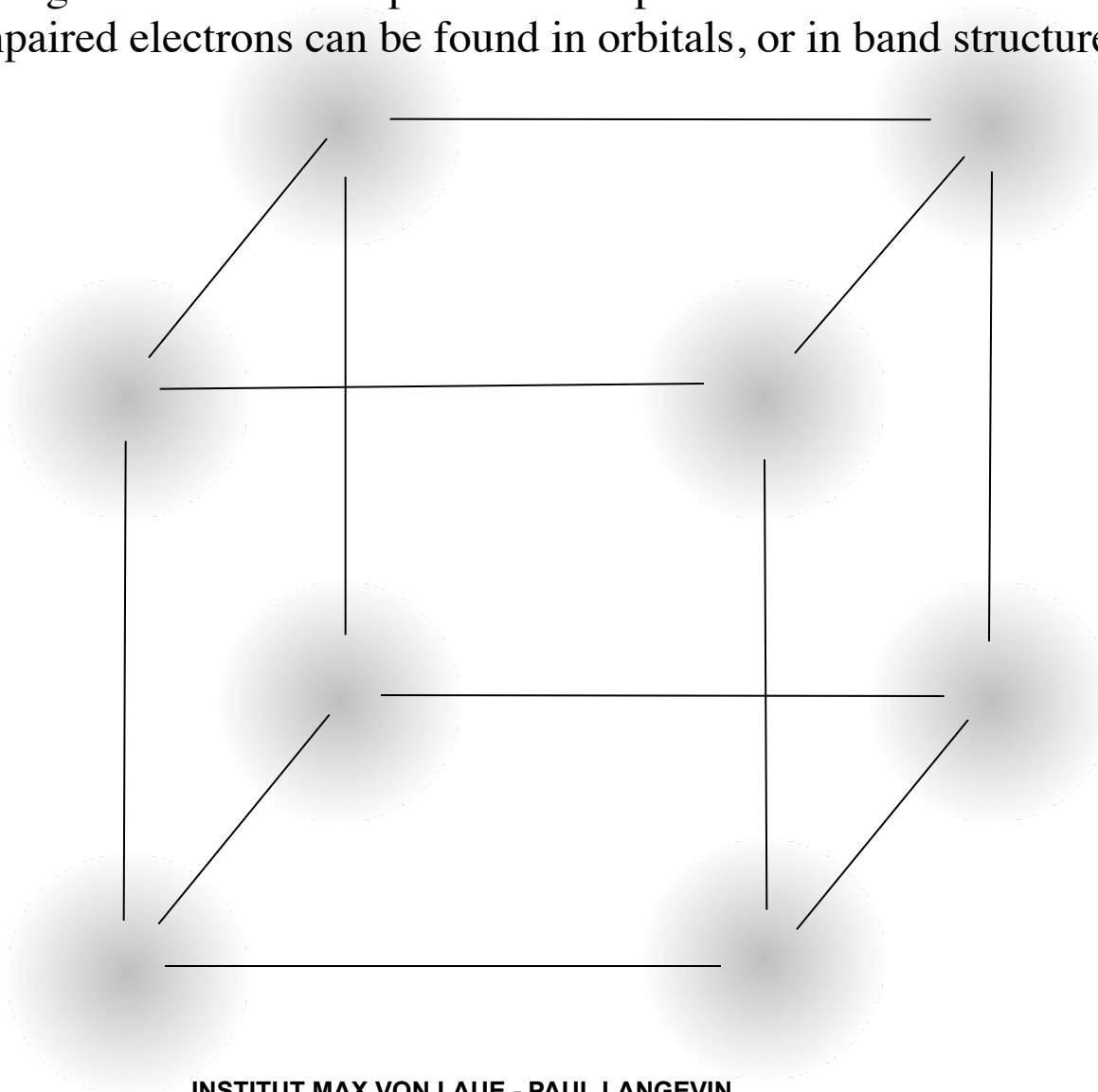
Unpaired electrons can be found in orbitals, or in band structures



Magnetic form factors

A magnetic moment is spread out in space

Unpaired electrons can be found in orbitals, or in band structures



Magnetic form factors

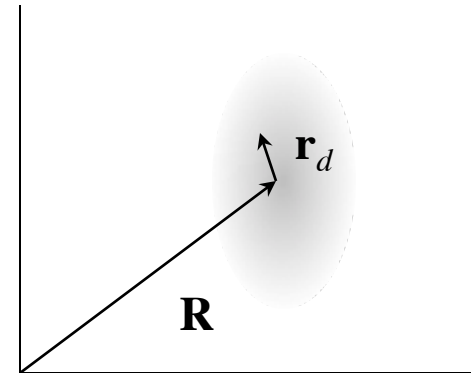
A magnetic moment is spread out in space

Unpaired electrons can be found in orbitals, or in band structures

e.g.

Take a magnetic ion with total spin $\mathbf{S}$ at position $\mathbf{R}$

The (normalized) density of the spin is $s_d(\mathbf{r})$
around the equilibrium position



$$\begin{aligned}
 \mathbf{M}(\mathbf{Q}) &= \int \mathbf{M}(\mathbf{r}) e^{i\mathbf{Q} \cdot \mathbf{r}} \cdot d\mathbf{r} \\
 &\propto \int s_d e^{i\mathbf{Q} \cdot \mathbf{r}_d} \cdot d\mathbf{r}_d \int S(\mathbf{R}) e^{i\mathbf{Q} \cdot \mathbf{R}} \cdot d\mathbf{R} \\
 &= f(Q) \int S(\mathbf{R}) e^{i\mathbf{Q} \cdot \mathbf{R}} \cdot d\mathbf{R} \\
 f(Q) &= \int s_d e^{i\mathbf{Q} \cdot \mathbf{r}_d} \cdot d\mathbf{r}_d
 \end{aligned}$$

$f(Q)$ is the *magnetic form factor*

It arises from the *spatial distribution of unpaired electrons* around a magnetic atom

Magnetic form factors

$f(Q)$ is the *magnetic form factor*

It arises from the *spatial distribution of unpaired electrons* around a magnetic atom

$$\frac{d\sigma_{\text{magnetic}}}{d\Omega} = \langle \mathbf{M}_{\perp}^*(\mathbf{Q}) \rangle \langle \mathbf{M}_{\perp}(\mathbf{Q}) \rangle$$

$$\propto f^2(Q) \int S_{\perp}(\mathbf{R}_i) S_{\perp}^*(\mathbf{R}_j) e^{i\mathbf{Q} \cdot (\mathbf{R}_i - \mathbf{R}_j)} \cdot d\mathbf{R}$$

There is no form factor for nuclear scattering, as the nucleus can be considered as a point compared to the neutron wavelength

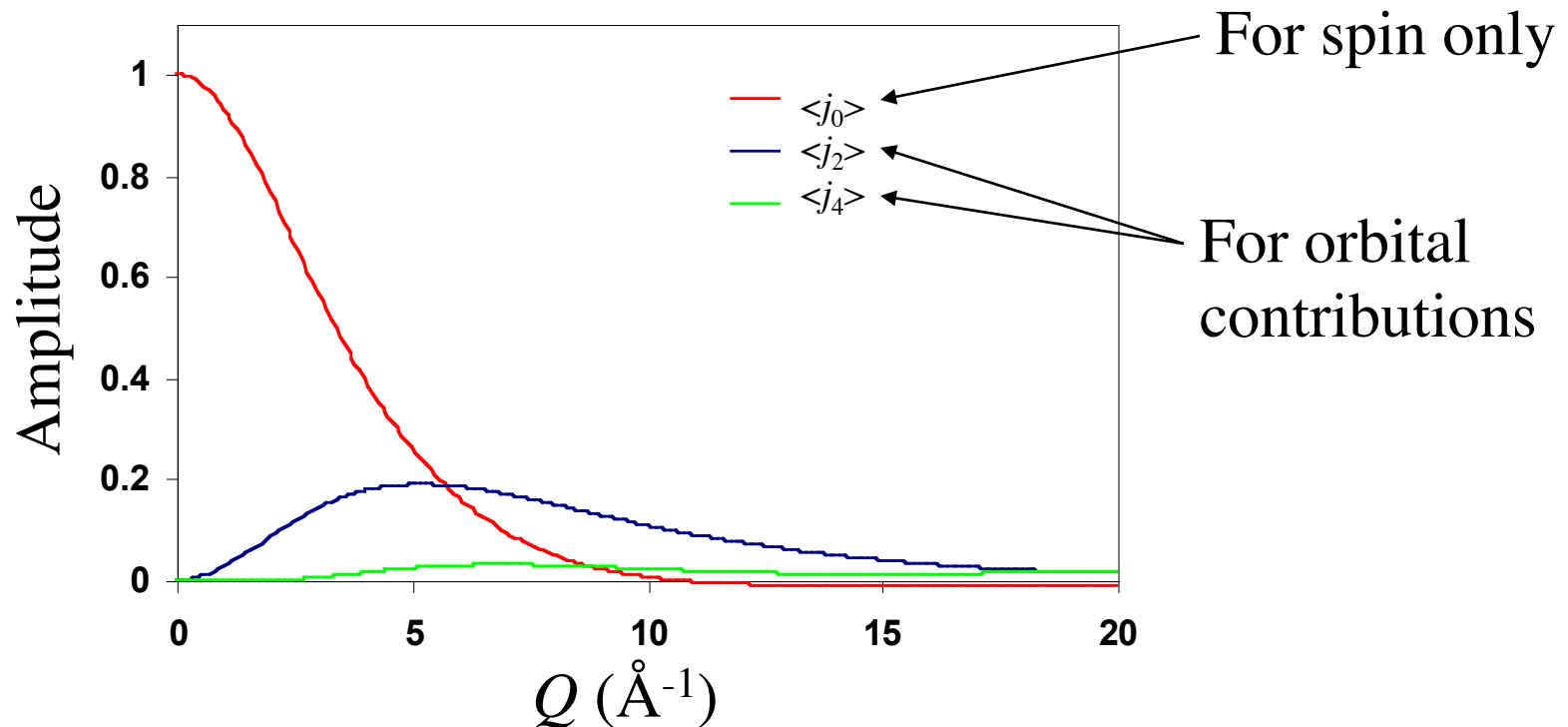
Magnetic form factors

Approximations for the form factors are tabulated

(P. J. Brown, International Tables of Crystallography, Volume C, section 4.4.5)

$$f(Q) = C_1 \langle j_0(Q/4\pi) \rangle + C_2 \langle j_2(Q/4\pi) \rangle + C_4 \langle j_4(Q/4\pi) \rangle + \dots$$

Form factors for iron



Nickel

H. A. Mook., Phys. Rev. **148** (1966) 495

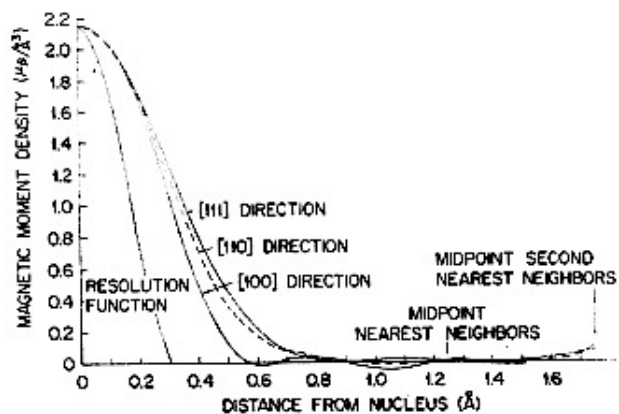


FIG. 1. Distribution of magnetic moment density along the three major crystallographic directions.

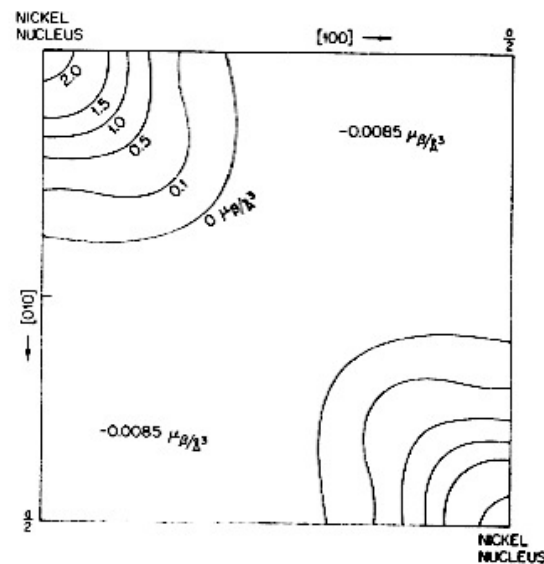


FIG. 4. The magnetic moment distribution in the [100] plane.

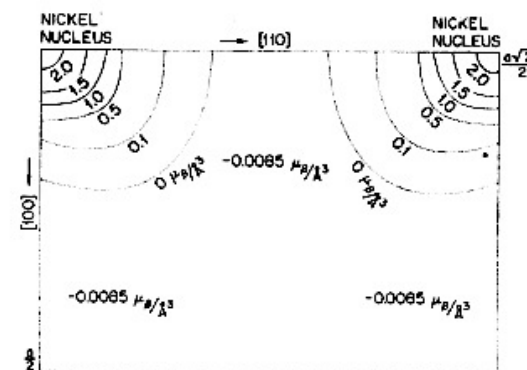
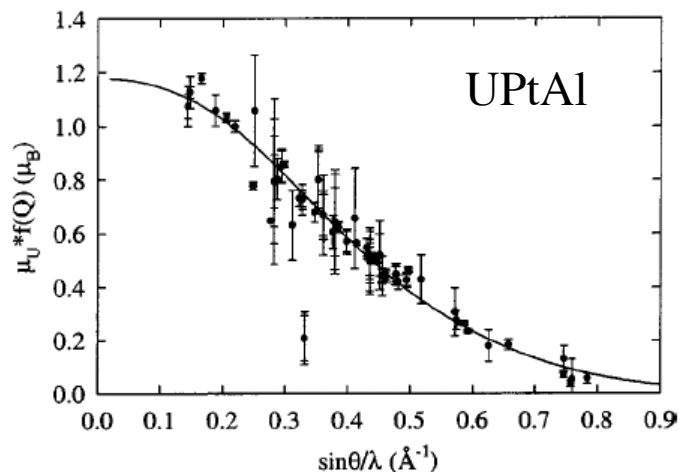


FIG. 5. The magnetic moment distribution in the [110] plane.

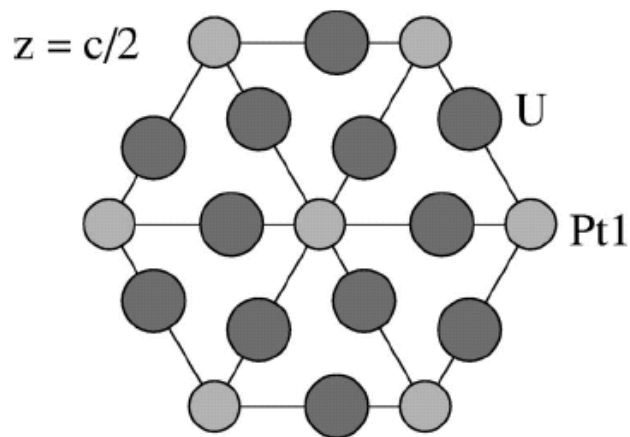
Magnetic electron density

$$f(Q) = \int s_d e^{iQ \cdot r_d} \cdot d\mathbf{r}_d$$

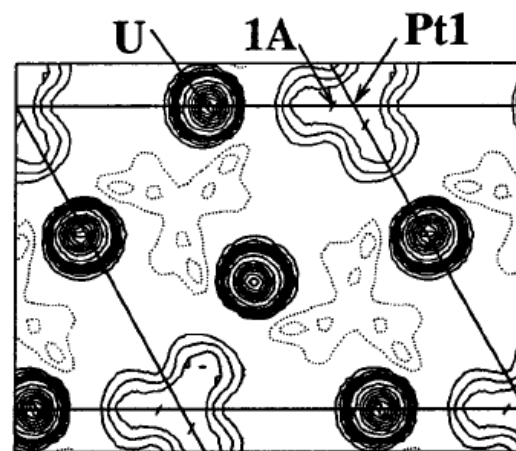


P. Javorsky *et al.*, Phys. Rev. B **67** (2003) 224429

From the form factors, the magnetic moment density in the unit cell can be derived



crystal structure



magnetic moment density

Magnetic fluctuations are governed by a wave equation:

$$H\psi = E \psi$$

The Hamiltonian is given by the physics of the material.

Given a Hamiltonian, H , the energies E can be calculated.
(this is sometimes very difficult)

Neutrons measure the energy, E , of the magnetic fluctuations, therefore probing the free parameters in the Hamiltonian.

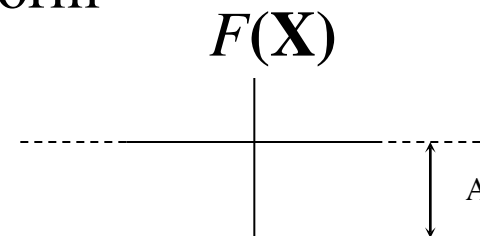
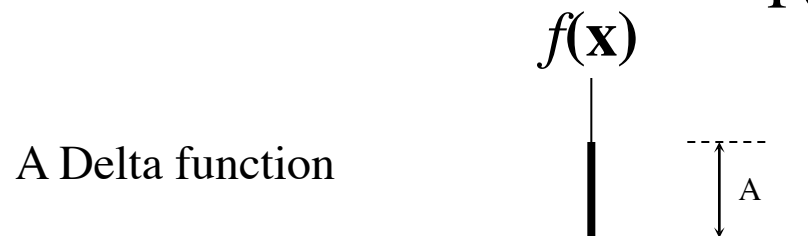
The crystal field is an electric field on an atom caused by neighbouring atoms in the sample.

It may lift the degeneracy of the energy levels for the atomic electrons.

If the electrons are unpaired, neutrons can cause the electrons to jump between the energy levels.

Because the crystal field is at a single atomic site, the inelastic scattering is essentially independent of $\mathbf{Q}$.

Fourier transform



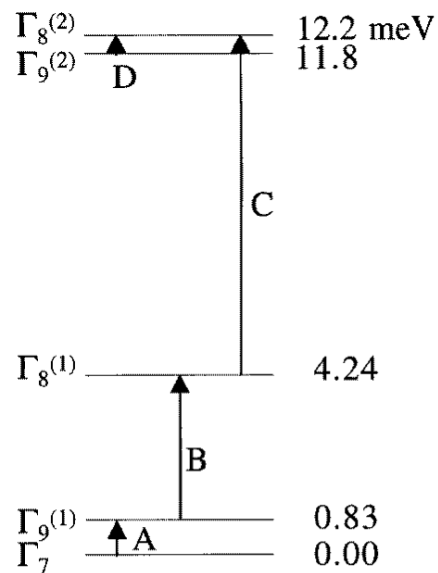
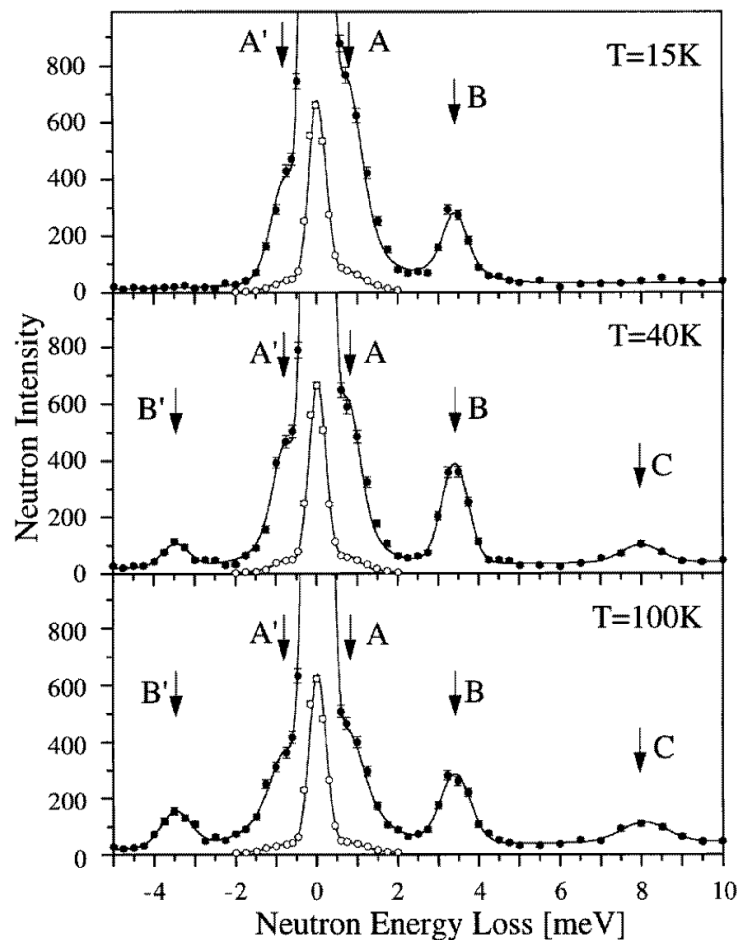
Crystal fields in NdPd_2Al_3

$$H = \sum_{l=2,4,6} B_l^0 O_l^0 + \sum_{l=2,4} B_l^4 O_l^4$$

O = Stevens parameters

(K. W. Stevens, Proc. Phys. Soc A **65** (1952) 209)

B = CF parameters,
measured by neutrons

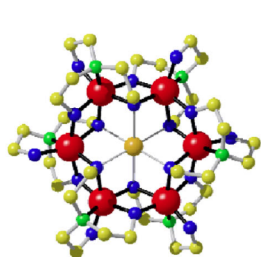


A. Dönni *et al.*, J. Phys.: Condens. Matter **9** (1997) 5921

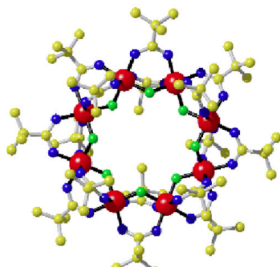
O. Moze., *Handbook of magnetic materials* vol. 11, 1998 Elsevier, Amsterdam, p.493

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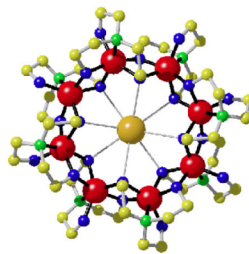
Molecular magnets



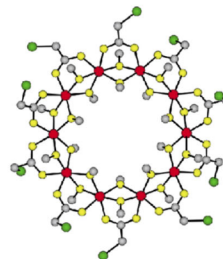
NaFe₆



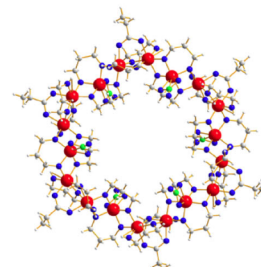
Cr₈



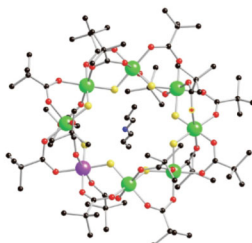
CsFe₈



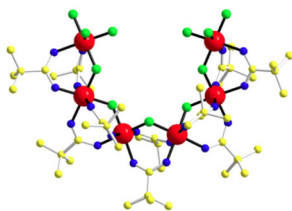
Fe₁₀



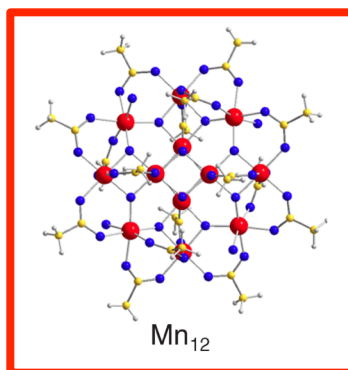
Fe₁₈



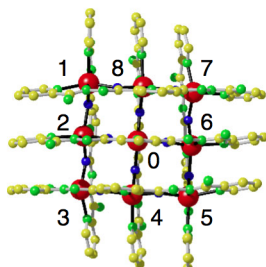
Cr₇Ni



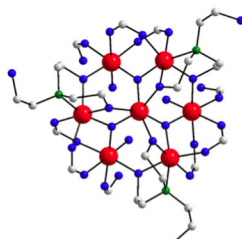
Cr₆



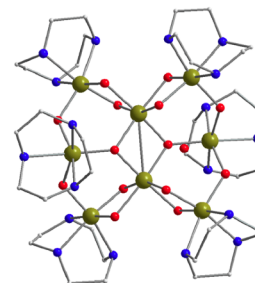
Mn₁₂



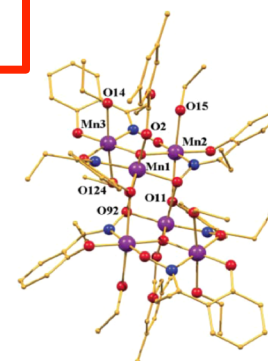
Mn-[3x3]



Fe₇



Fe₈

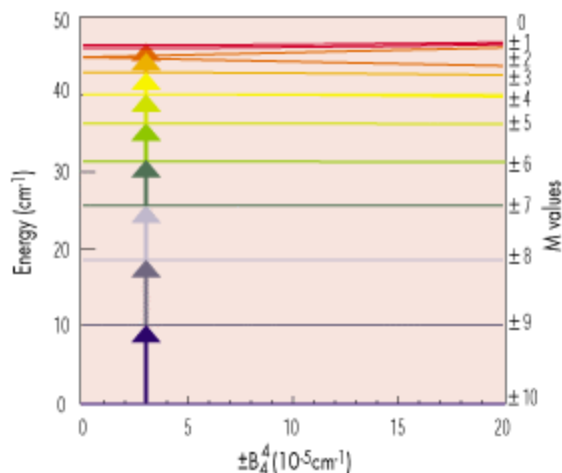


Mn₆

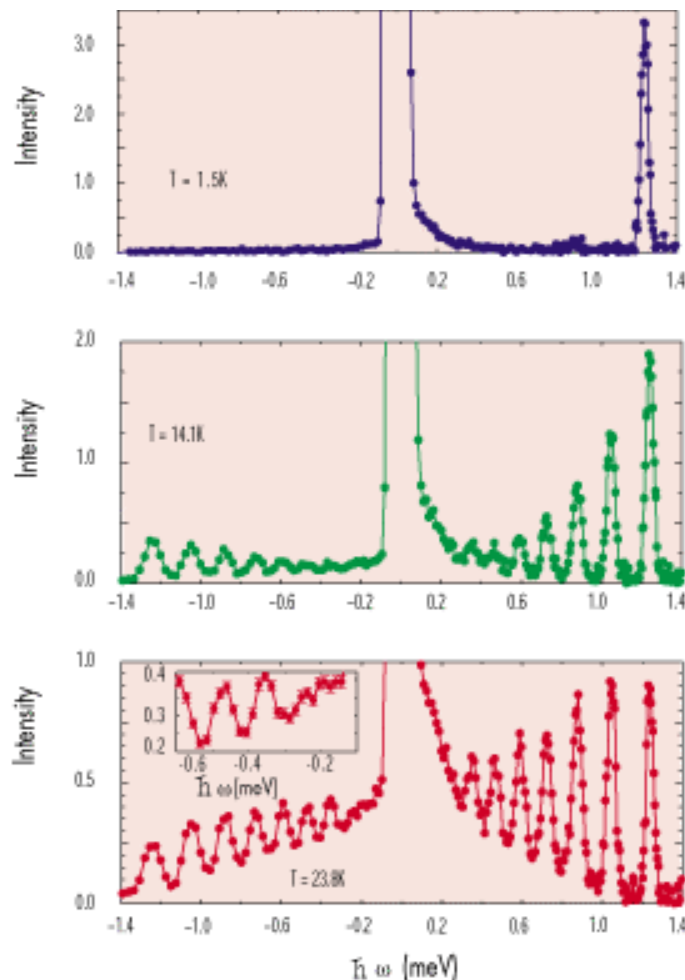
A. Furrer and O. Waldmann, Rev. Mod. Phys. **85** (2013) 367

Quantum tunneling in Mn₁₂-acetate

$$H = \sum_{l=2,4,6} B_l^0 O_l^0 + \sum_{l=2,4} B_l^4 O_l^4$$

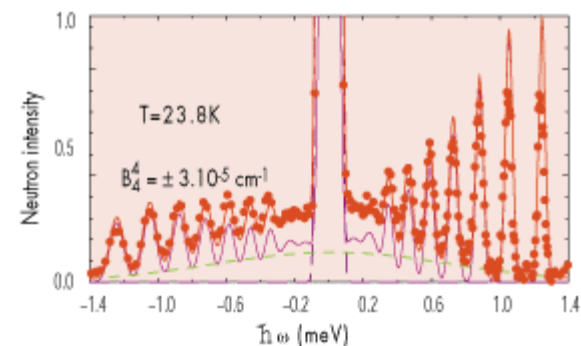


Calculated energy terms



Neutron spectra

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Fitted data with scattering from:

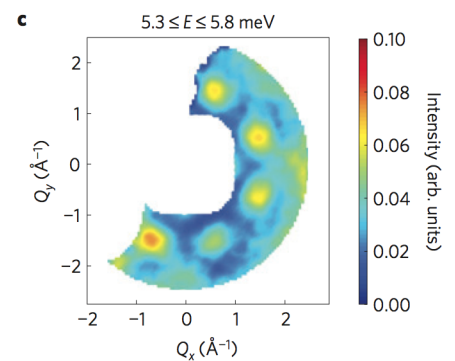
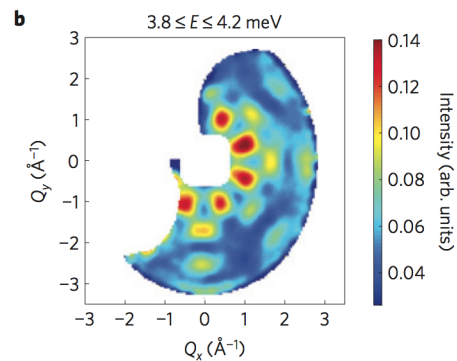
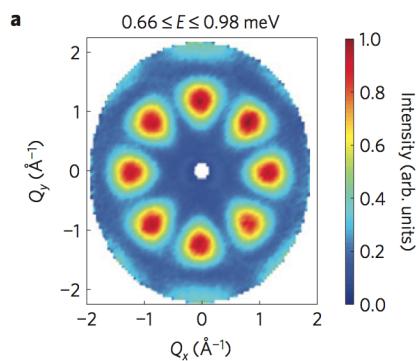
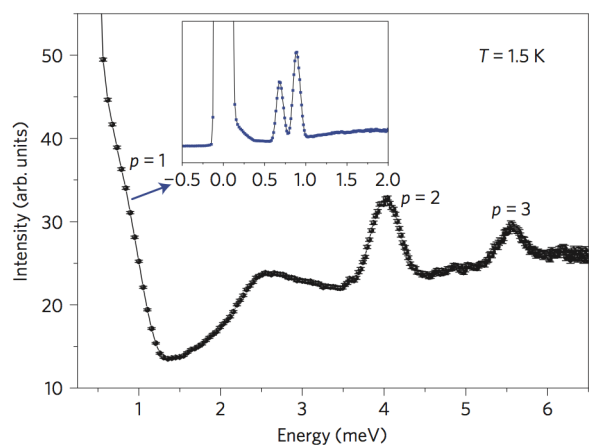
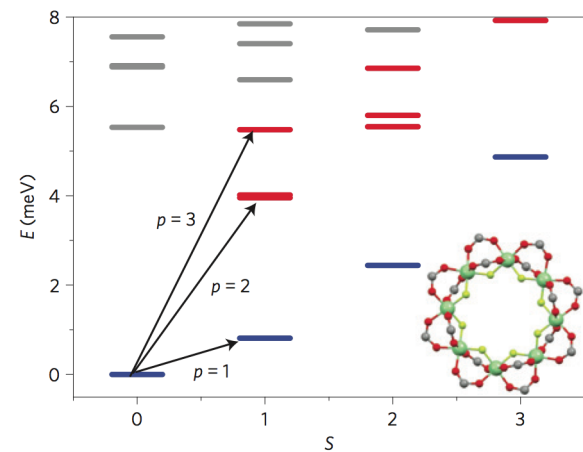
- energy levels
- elastic scattering
- incoherent background

I. Mirebeau *et al.*, Phys. Rev. Lett. **83** (1999) 628

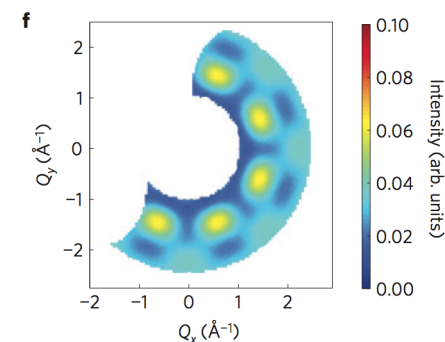
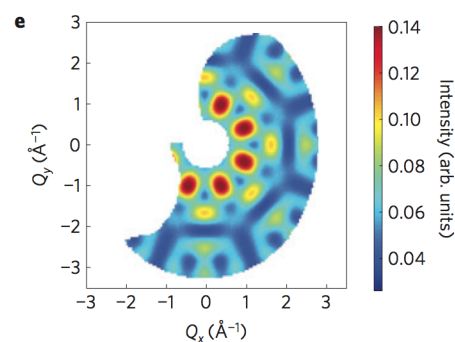
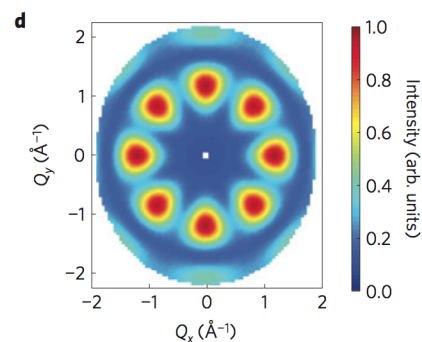
R. Bircher *et al.*, Phys. Rev. B **70** (2004) 212413

Space-time correlations in Cr₈

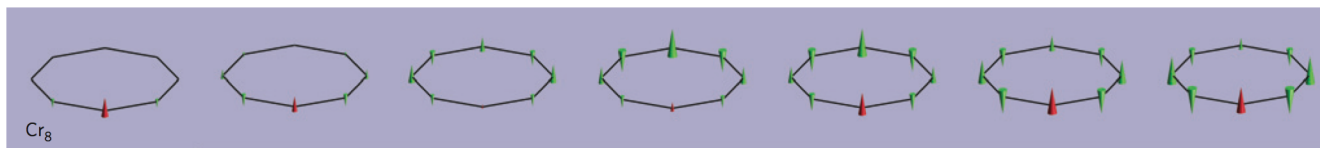
$$H = J \sum_{i=1}^8 \mathbf{s}_i \cdot \mathbf{s}_{i+1}$$



Data



Fit



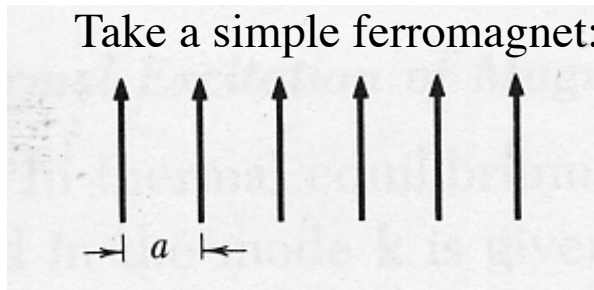
Spin waves and magnons

A simple Hamiltonian for spin waves is:

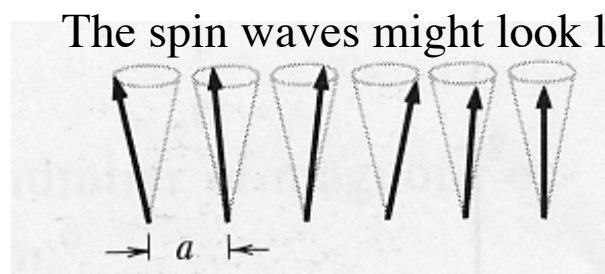
$$H = -J \sum_{i,j} \mathbf{s}_i \mathbf{s}_j$$

J is the magnetic exchange integral, which can be measured with neutrons.

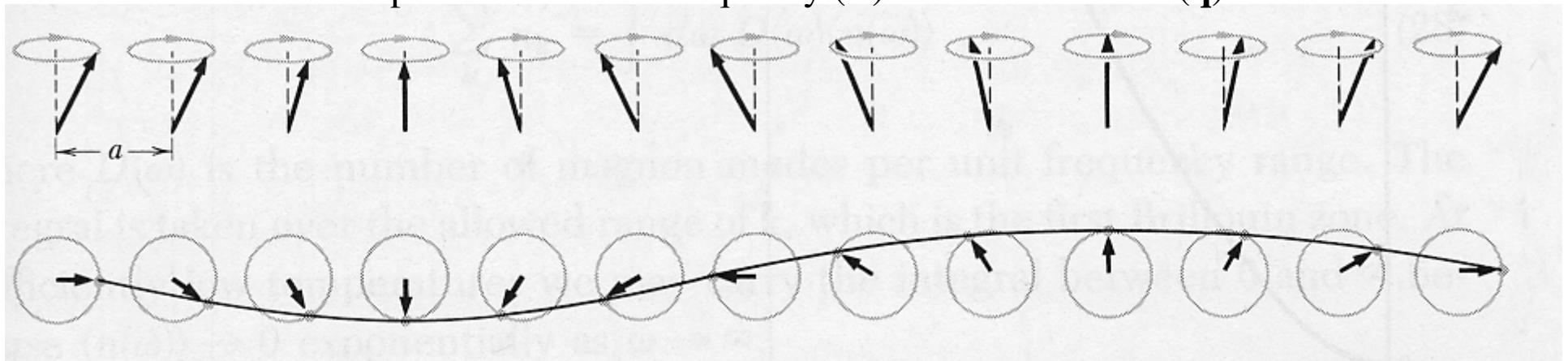
Take a simple ferromagnet:



The spin waves might look like this:

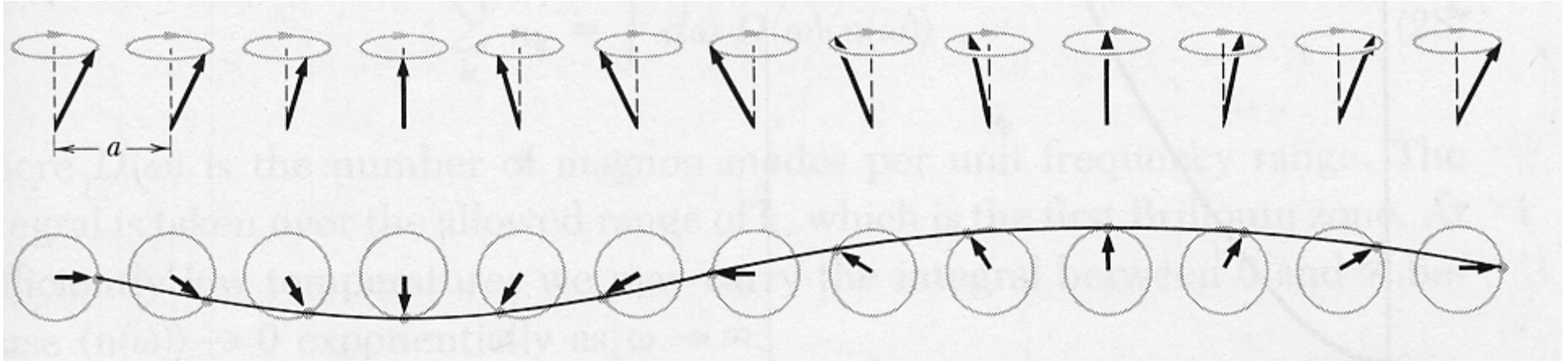


Spin waves have a frequency (ω) and a wavevector ($\mathbf{q}$)

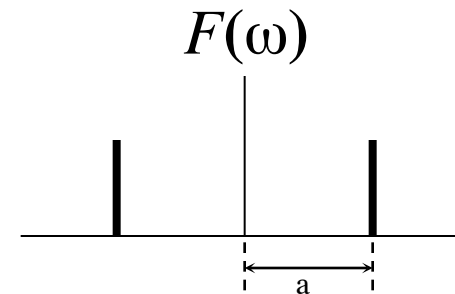
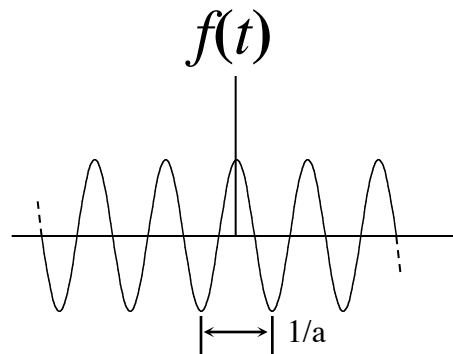


The frequency and wavevector of the waves are *directly measurable* with neutrons

Magnons



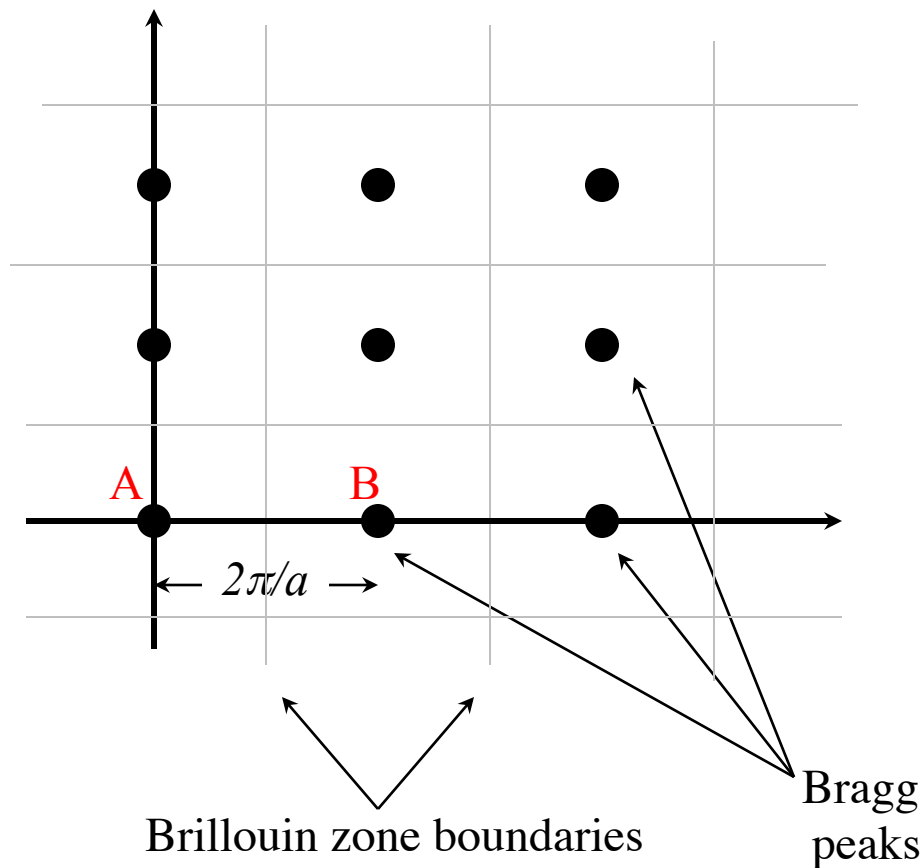
The Fourier Transform for a periodic function:



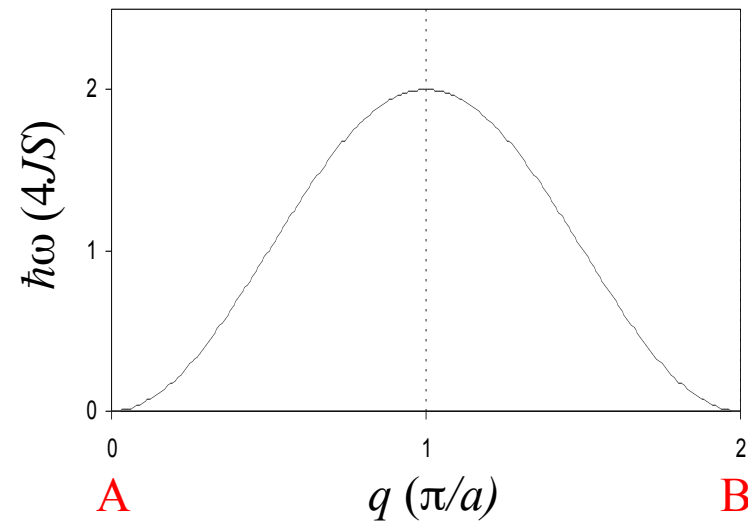
Each wavelength for the magnon has its own periodicity.
 Each wavevector for the magnon has its own frequency (energy)

Magnons and reciprocal space

Reciprocal space



Spin wave dispersion

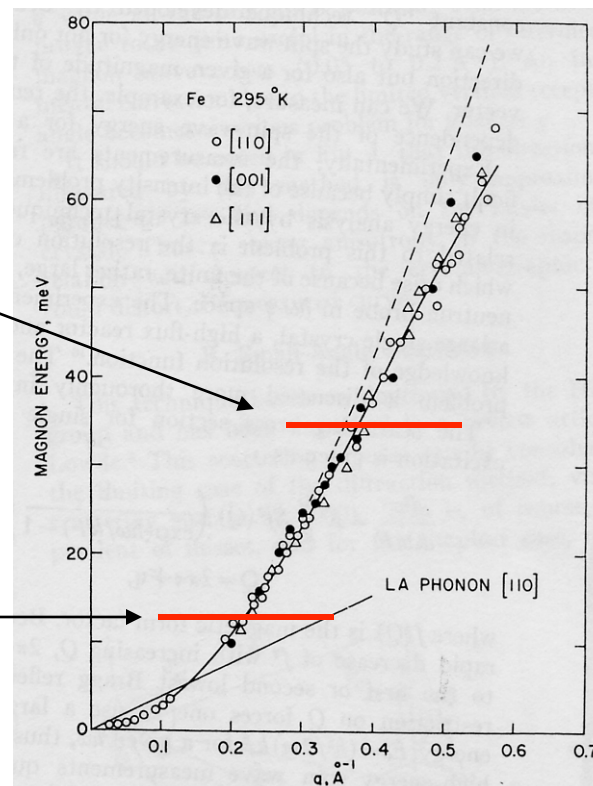
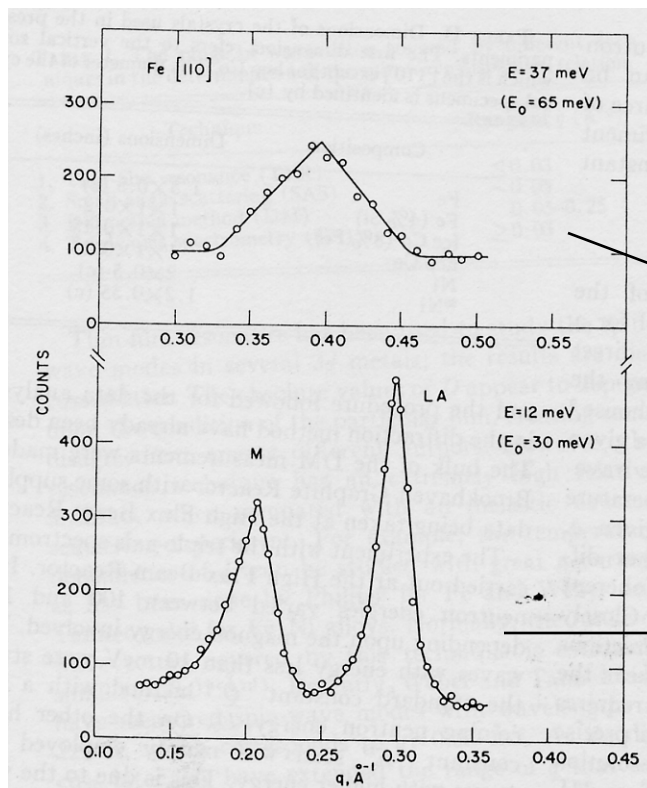
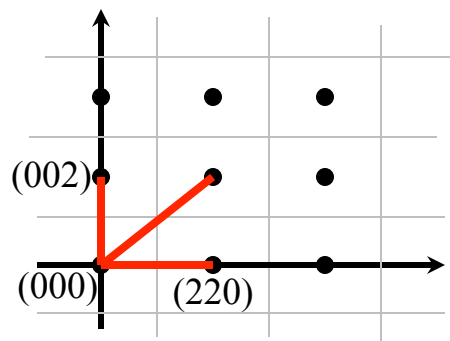


$$\begin{aligned}\hbar\omega &= 4JS(1 - \cos qa) \\ &= Dq^2 \quad (\text{for } qa \ll 1) \\ D &= 2JSa^2\end{aligned}$$

C. Kittel, *Introduction to Solid State Physics*, 1996, Wiley, New York

F. Keffer, *Handbuch der Physik* vol 18II, 1966 Springer-Verlag, Berlin

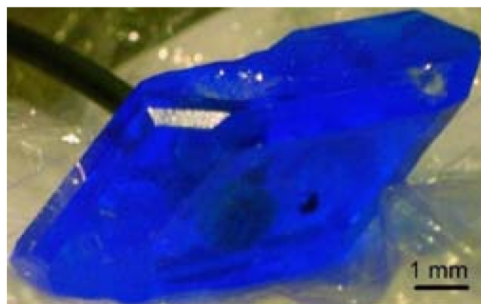
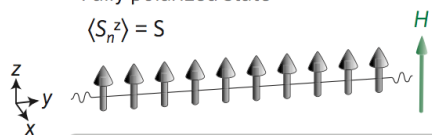
Magnons in crystalline iron



G. Shirane *et al.*, J. Appl. Phys. **39** (1968) 383

Magnetic excitations in CuSO_4

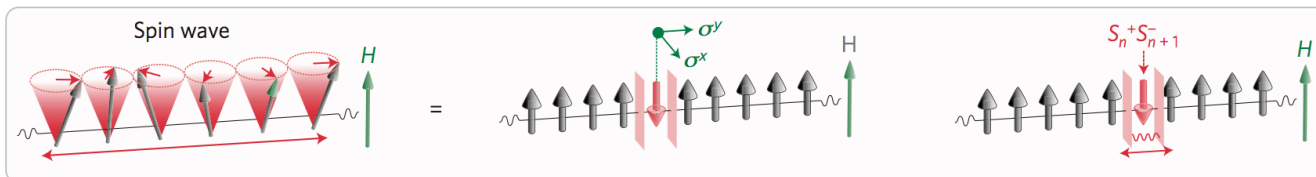
Fully polarized state
 $\langle S_n^z \rangle = S$



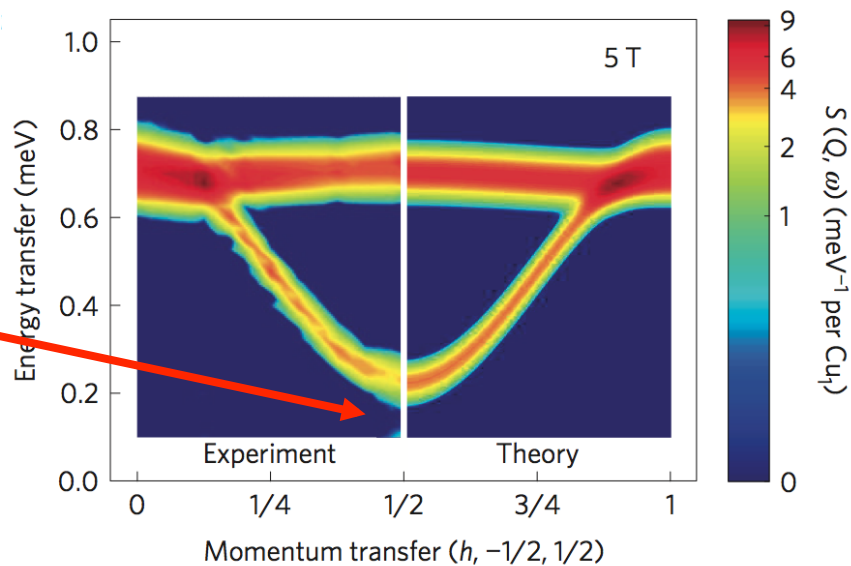
Time 0



Time t



Gap in dispersion due to magnetic anisotropy

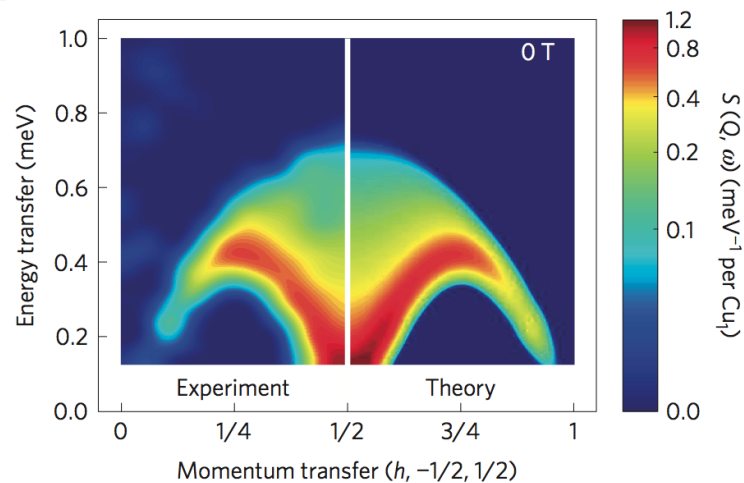
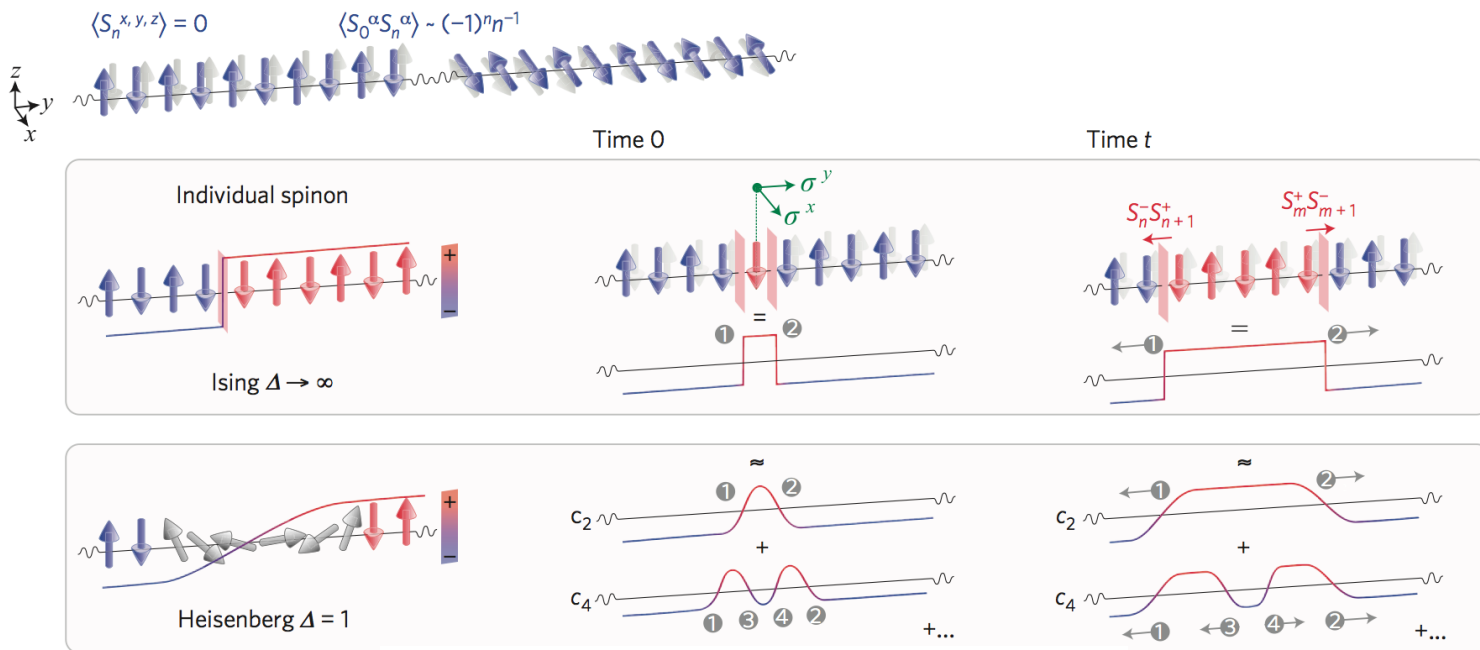


M. Mourigal *et al.*,

Nature Phys. **9** (2013) 435

Magnetic excitations in CuSO_4

Zero magnetic field state



M. Mourigal *et al.*,

Nature Phys. **9** (2013) 435

Neutron scattering and magnetic susceptibility

Magnetic susceptibility is a fundamental property of a material. It is defined as:

$$\chi = \frac{M}{H}$$

In a magnetic system, $\mathbf{M}$ is a vector which varies as a function of space, $\mathbf{r}$, and (due to fluctuations) as a function of time, t .

The time is related to the susceptibility by:

$$M_{\alpha}(t) \propto \chi_{\alpha\alpha}(\omega) H_{0\alpha} e^{-i\omega t} + \chi_{\alpha\alpha}^*(\omega) H_{0\alpha}^* e^{i\omega t}$$

The susceptibility is a complex tensor:

$$\chi_{\alpha\alpha}(\omega) = \chi'_{\alpha\alpha}(\omega) + i\chi''_{\alpha\alpha}(\omega)$$

The rate of energy gain is given by:

$$\frac{d\bar{E}}{dt} = -M_{\alpha} \frac{dH}{dt} \propto \chi''_{\alpha\alpha}(\omega)$$

and the inelastic cross-section is then related to a *generalized* susceptibility:

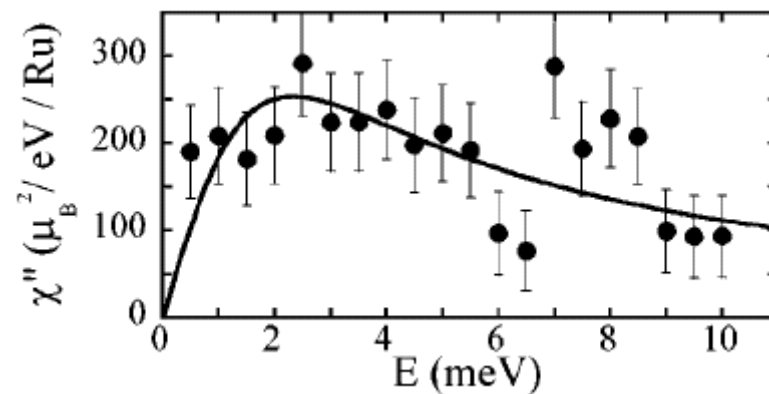
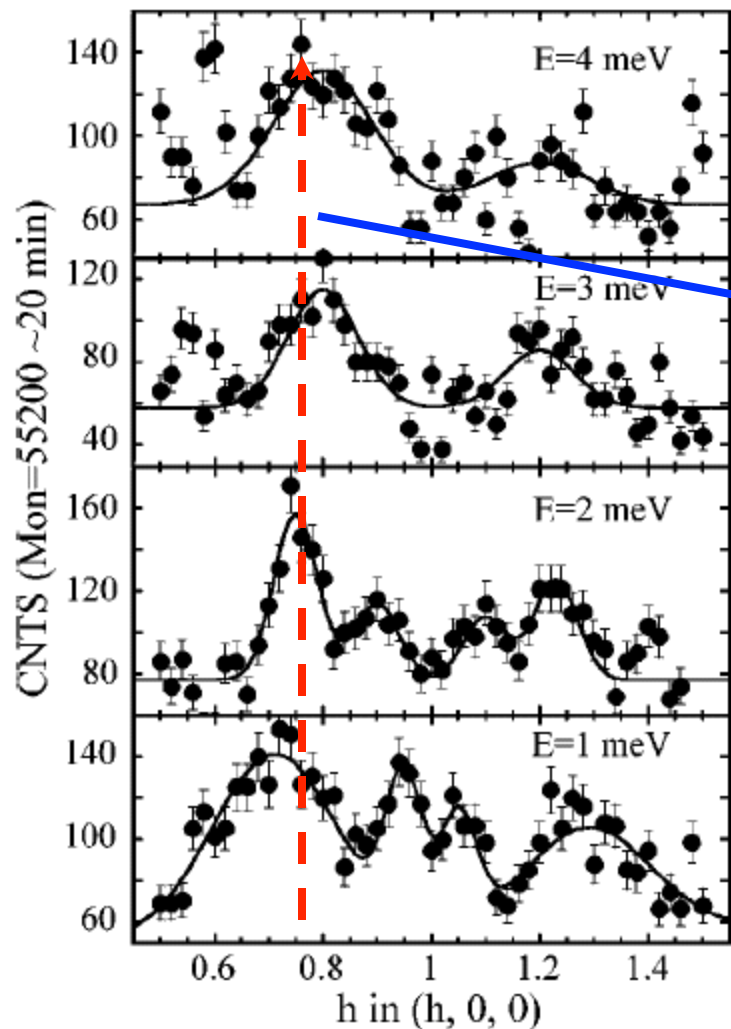
$$\frac{d^2\sigma}{d\Omega dE} \propto \frac{k'}{k} \sum_{\alpha} \frac{\chi''_{\alpha\alpha}(\mathbf{Q}, \omega)}{(1 - e^{-\beta\hbar\omega})}$$

T. J. Hicks, *Magnetism in disorder*, Oxford University Press, Oxford, 1995

S. W. Lovesey, *Theory of neutron scattering from condensed matter*, Oxford University Press, Oxford, 1986

Spin excitations in $\text{Sr}_3\text{Ru}_2\text{O}_7$

$\text{Sr}_3\text{Ru}_2\text{O}_7$ is from a family of materials that are low-dimensional magnetic, and superconductors

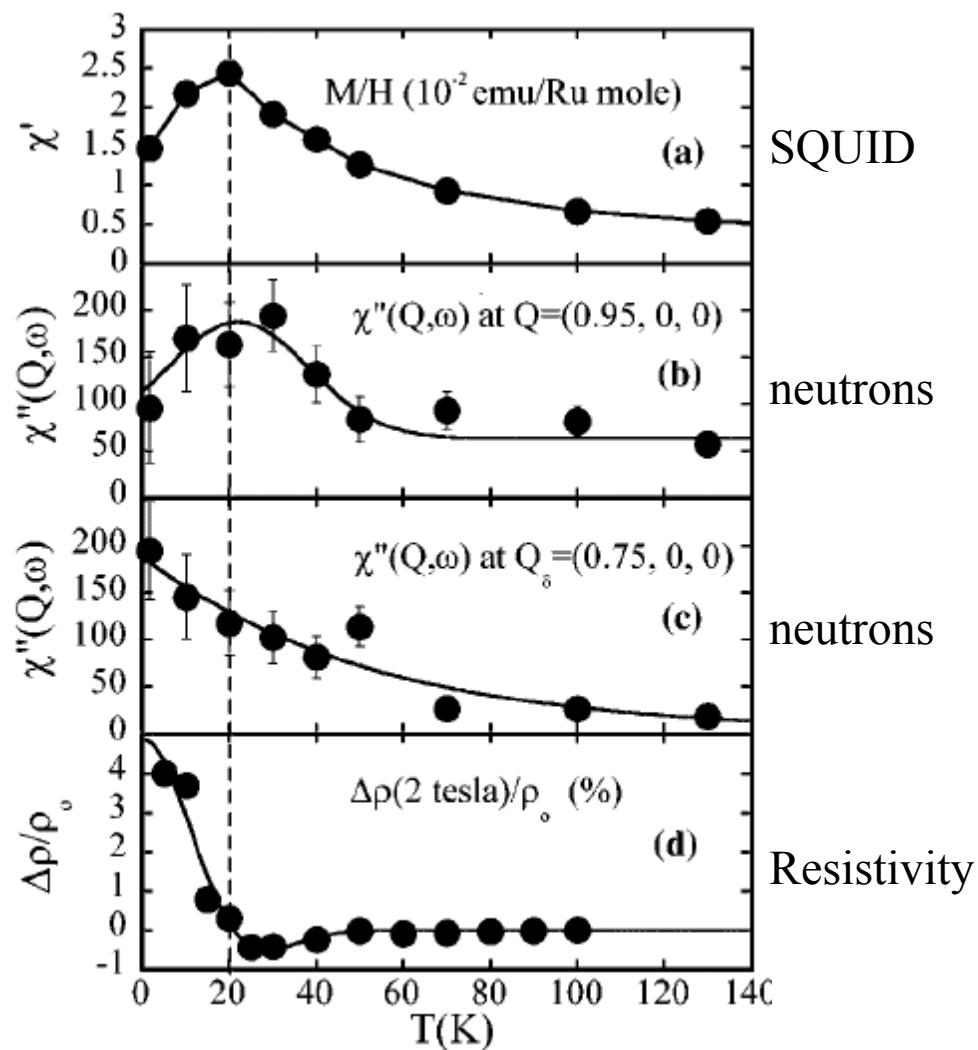
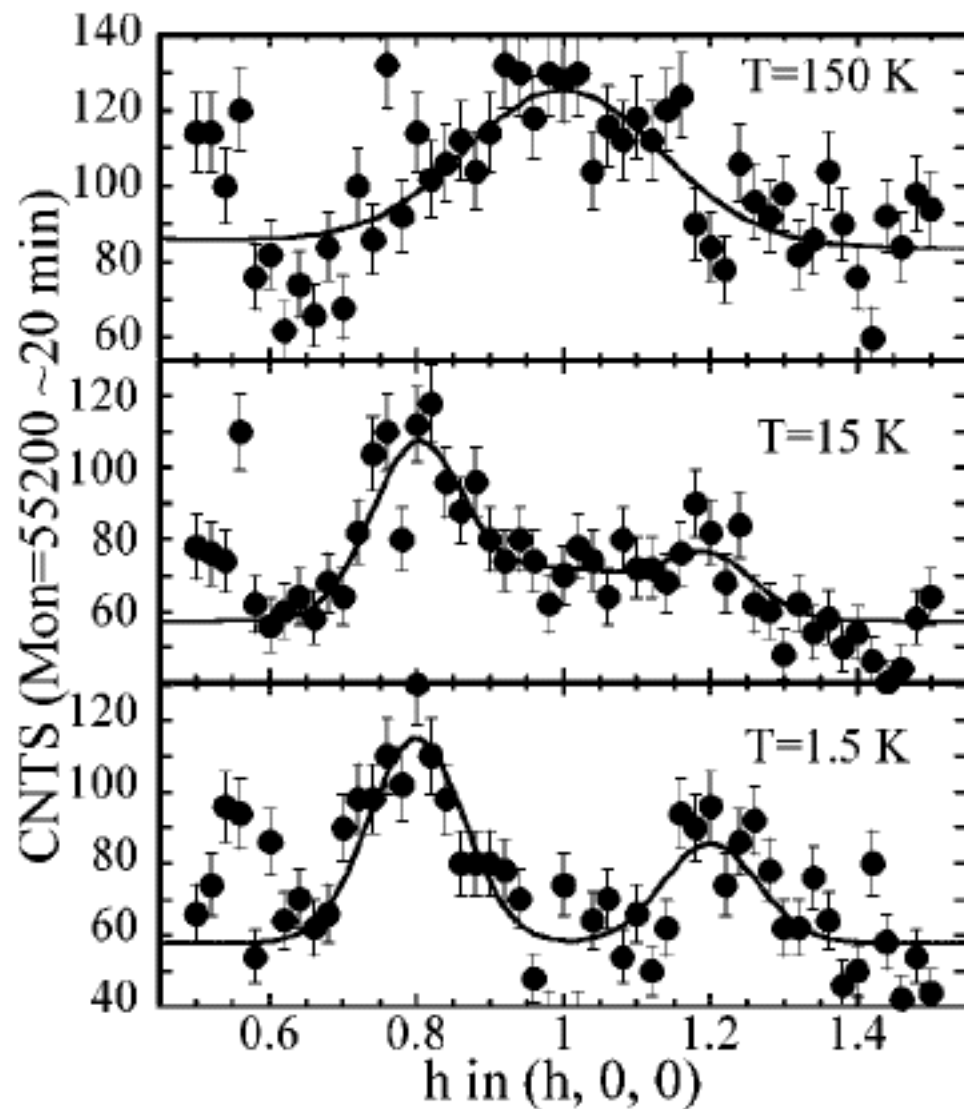


Susceptibility

L. Capogna *et al.*, Phys. Rev. B. **87** (1998) 143

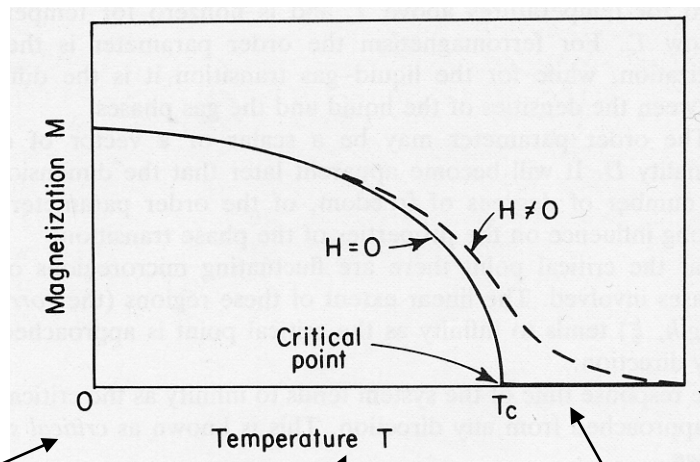
Inelastic neutron scattering at 1.5K

Temperature dependence of the spin excitations in $\text{Sr}_3\text{Ru}_2\text{O}_7$

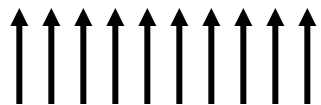


Magnetic phase transitions

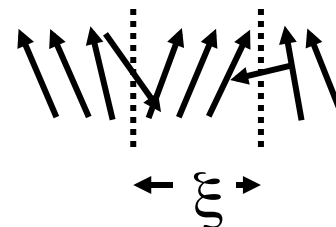
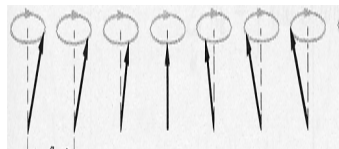
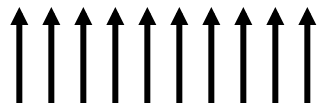
The magnetic structure of a simple ferromagnet as a function of temperature



Time-average
($\hbar\omega = 0$)



Snapshot



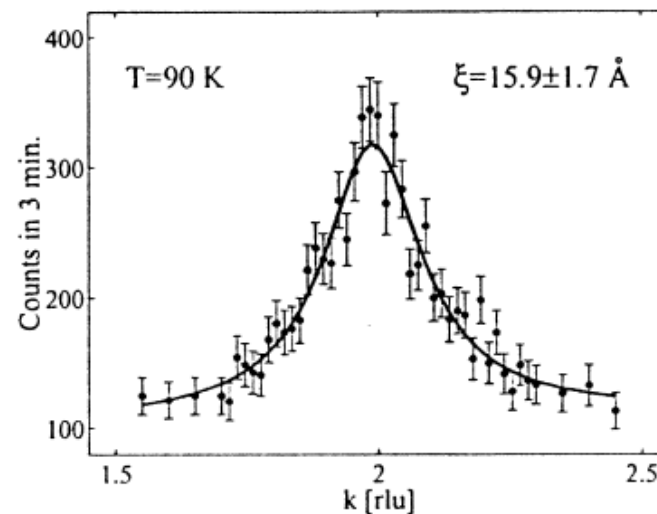
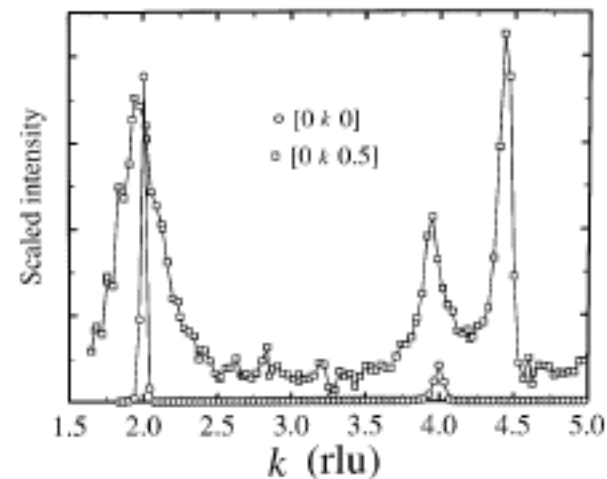
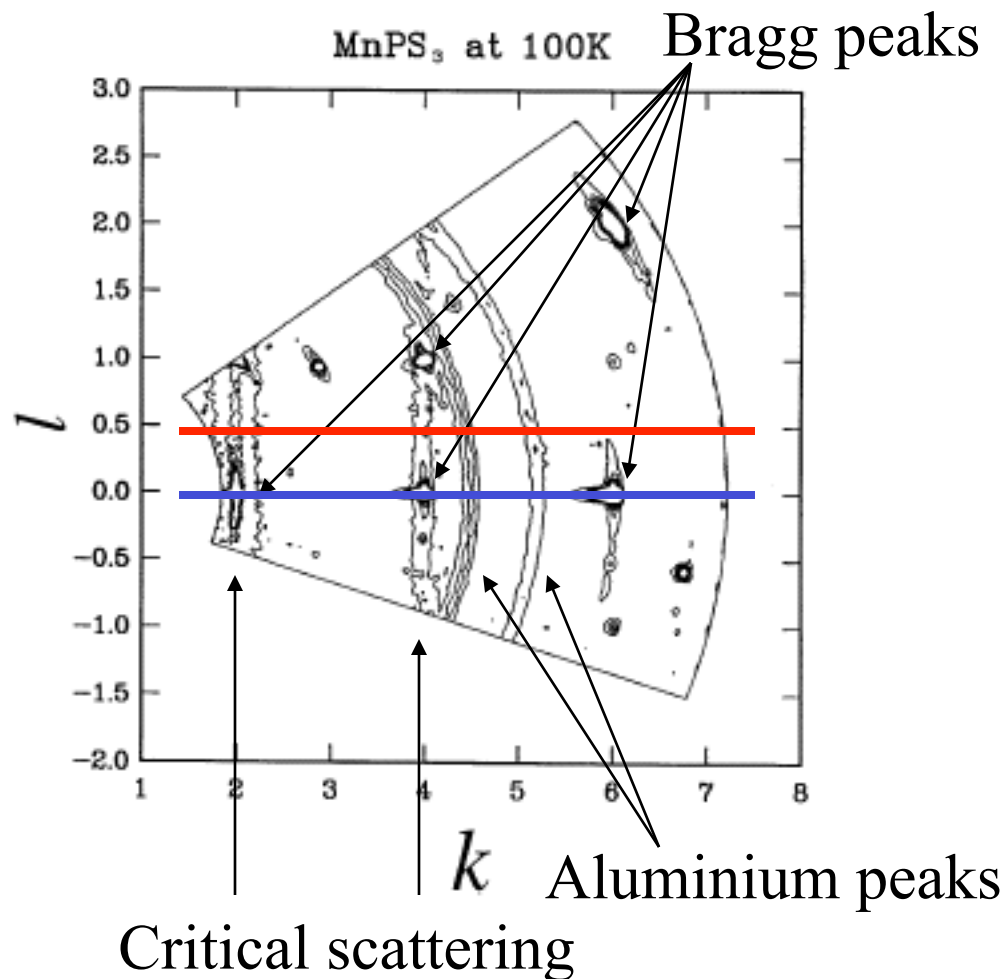
$(\int_{-\infty}^{\infty} \hbar\omega)$

Renormalization theory and universality:

$$M(T) \propto (T-T_N)^{-\beta} \quad \xi(T) \propto (T-T_N)^{-\nu}$$

M. F. Collins, *Magnetic critical scattering*, 1989, Oxford University, Oxford

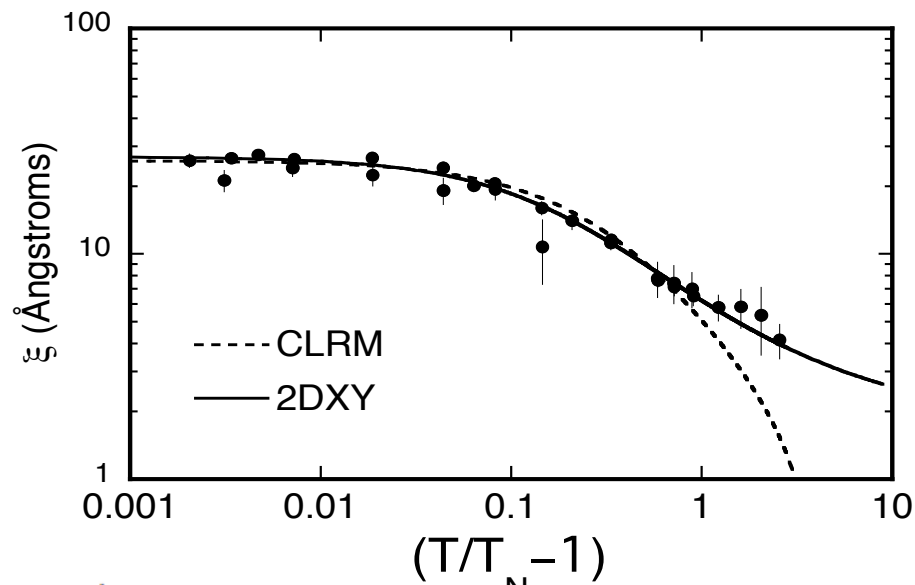
Magnetic critical scattering in MnPS_3



A.R. Wildes *et al.*, J. Magn. Magn. Mater. **87** (1998) 143

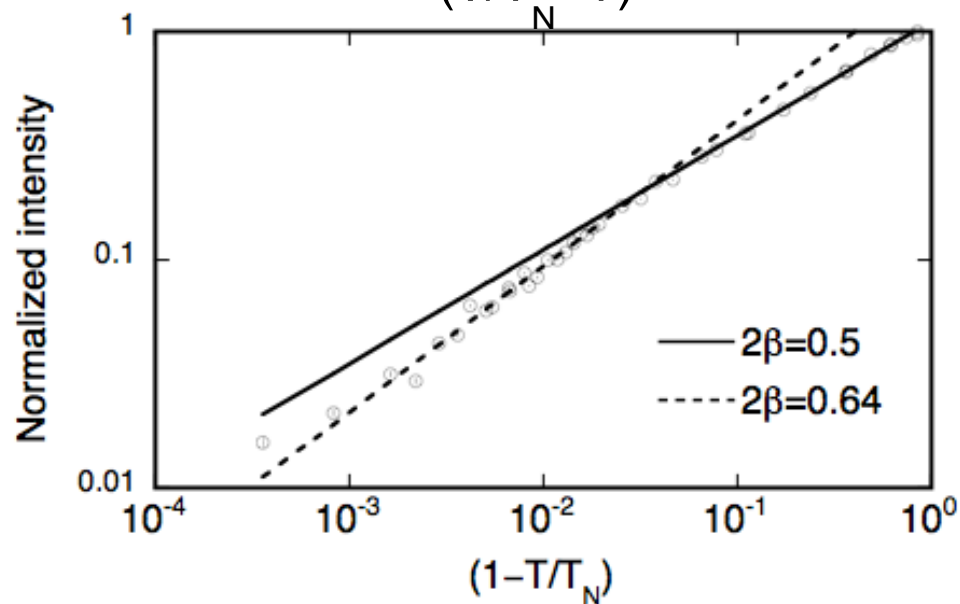
Correlation length (from the widths of the rods, *energy integrated*)

H.M. Rønnow *et al.*, Physica B **276-278** (2000) 676



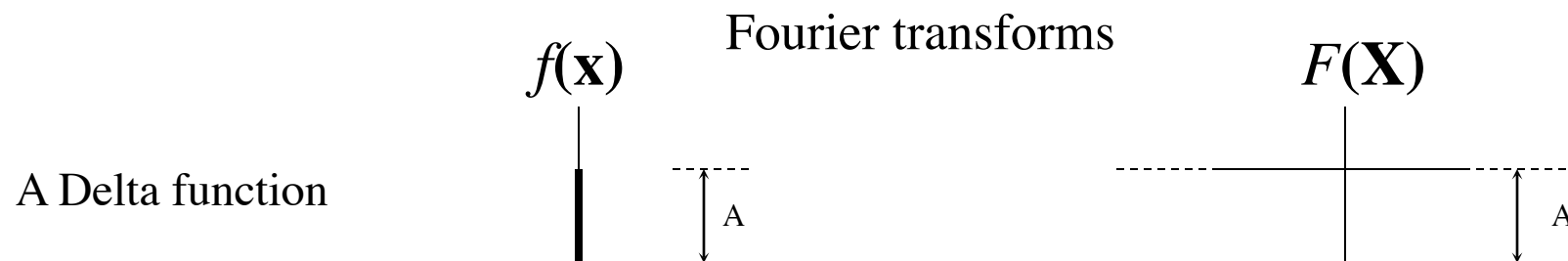
Sublattice magnetization (from the intensities of the Bragg peaks, *zero energy transfer*)

A. R. Wildes *et al.*, PRB **74** (2006) 094422



Diffuse scattering from a paramagnet

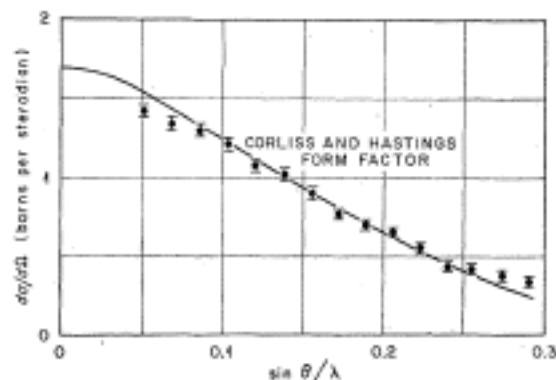
An ideal paramagnet has no correlation in space *or* time



For a paramagnet, $z(\mathbf{r}) = \delta(r)$, and must *integrate* over energy

$$\frac{d\sigma^{\pm\mp}}{d\Omega} \propto \frac{2}{3} f^2(Q) S(S+1)$$

Paramagnetic scattering from MnF_2



R. M. Moon, T. Riste and W. K. Koehler, Phys. Rev. **181** (1969) 920

Neutrons and bulk susceptibility

If neutrons measure the generalized susceptibility,
it must be possible to convert between bulk susceptibility measurements and neutron cross-sections.
This can be done using the Kramers-Krönig relation:

$$\int d\omega \frac{\chi''(\mathbf{Q}, \omega)}{\omega} = \pi \chi'(\mathbf{Q}, \omega)$$

Integrate the neutron scattering over all energies:

$$\begin{aligned} \frac{d\sigma}{d\Omega} &= \int \hbar d\omega \left(\frac{d^2\sigma}{d\Omega d\omega} \right) \\ &\propto k_B T \sum_{\alpha} \chi'_{\alpha\alpha}(\mathbf{Q}, 0) \end{aligned}$$

Bulk susceptibility measures the real part of the susceptibility.
Bulk susceptibility averages over all the sample, which is equivalent to $\mathbf{Q} = 0$.

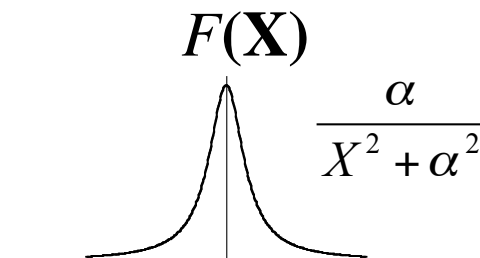
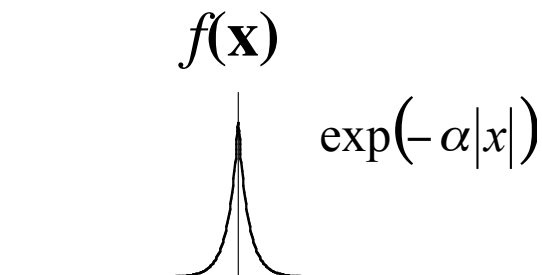
$$i.e. \frac{d\sigma}{d\Omega}(\mathbf{Q} = 0) \propto k_B T \chi'$$

Bulk susceptibility can be put as a point on a neutron scattering plot!

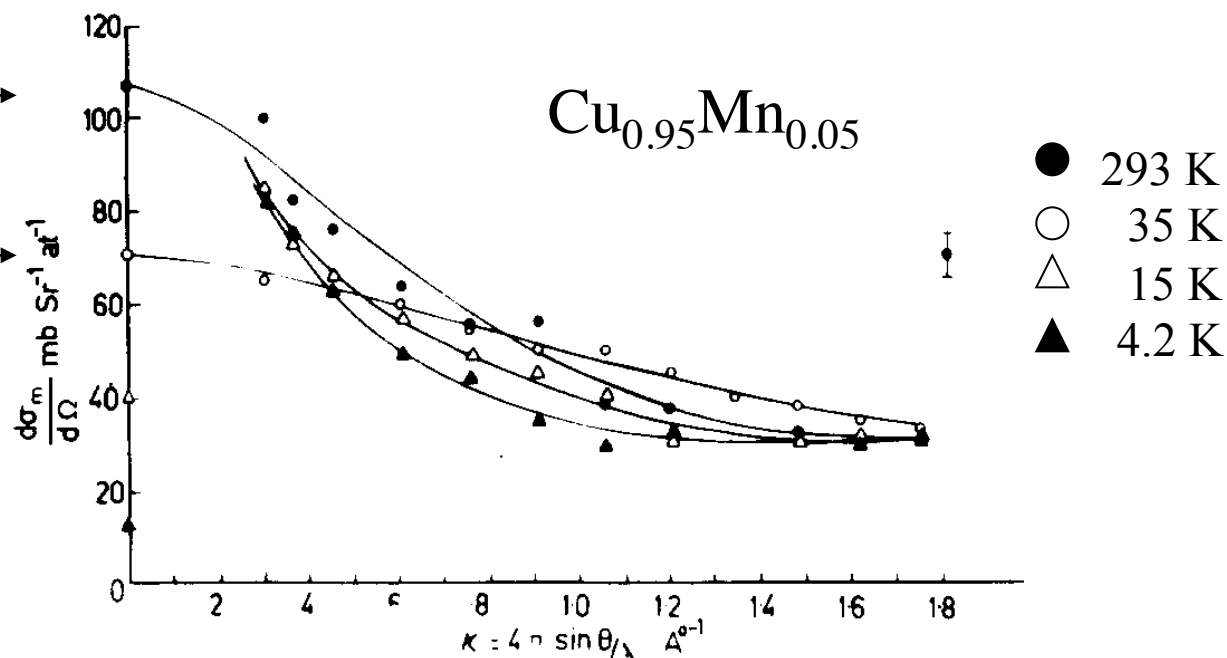
Diffuse scattering from a paramagnet

Paramagnets/spin glasses always have some correlations, particularly in time

An exponential



Points taken
from bulk
susceptibility



N. Ahmed and T. J. Hicks, Solid State Comm. **15** (1974) 415

T. J. Hicks, *Magnetism in disorder*, 1995, Clarendon, Oxford

Diffuse scattering and short-ranged order

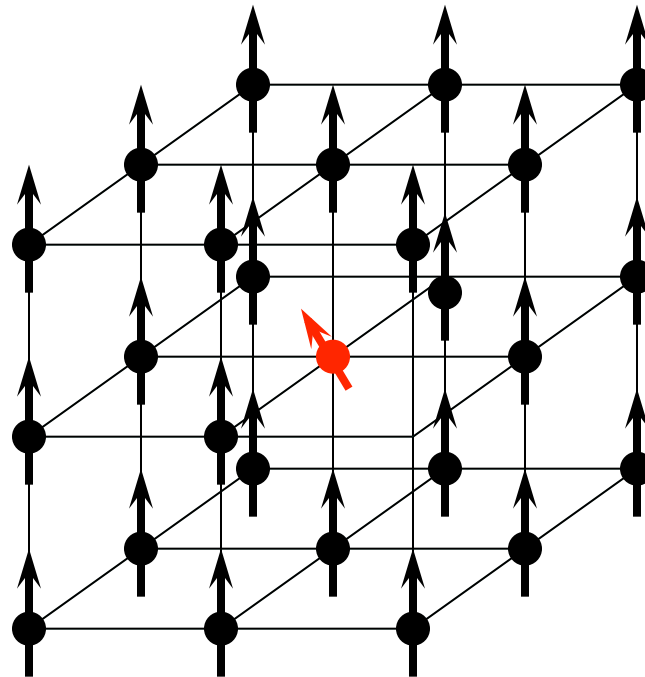
$$\left(\left| \langle \hat{V}^2 \rangle \right| - \left| \langle \hat{V} \rangle \right|^2 \right) \int z(\mathbf{r}) e^{i\mathbf{Q} \cdot \mathbf{r}} \cdot d\mathbf{r}$$

Take a magnetic moment at the origin.

$z(\mathbf{r})$ describes the probability of finding the same moment at position $\mathbf{r}$.

In a crystalline material the integral becomes a sum over neighbours

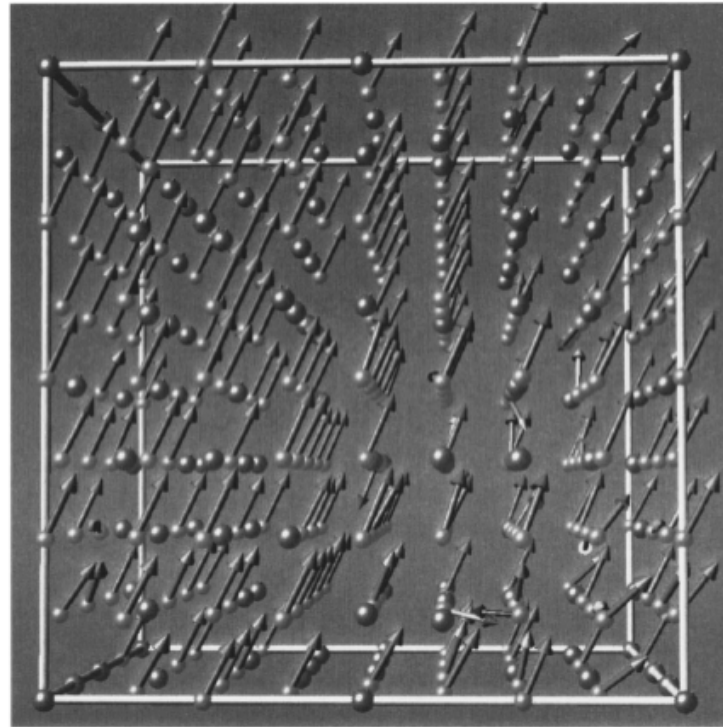
$$\propto S(S+1) + \sum_{R_n} \langle \mathbf{S}_0 \cdot \mathbf{S}_{R_n} \rangle \frac{\sin(QR_n)}{QR_n}$$



Diffuse scattering and short-ranged order

“Short-ranged” order can still extend over many nanometres

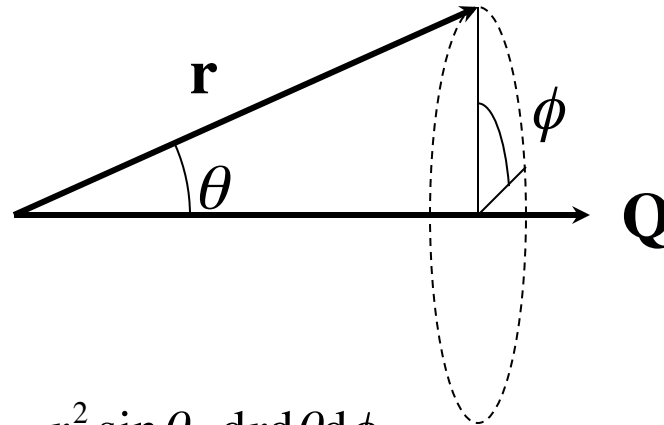
Calculation of the magnetic structure of $\text{Ni}_{0.65}\text{Fe}_{0.35}$



256 atoms, Yang *et al.* J. Appl. Phys. **81** (1997) 3973

Isotropic Fourier Transforms

Most magnetic diffuse scattering experiments are done on powders
In the case of scattering that is isotropic in three dimensions:



$$d\mathbf{r} = r^2 \sin\theta \cdot dr d\theta d\phi$$

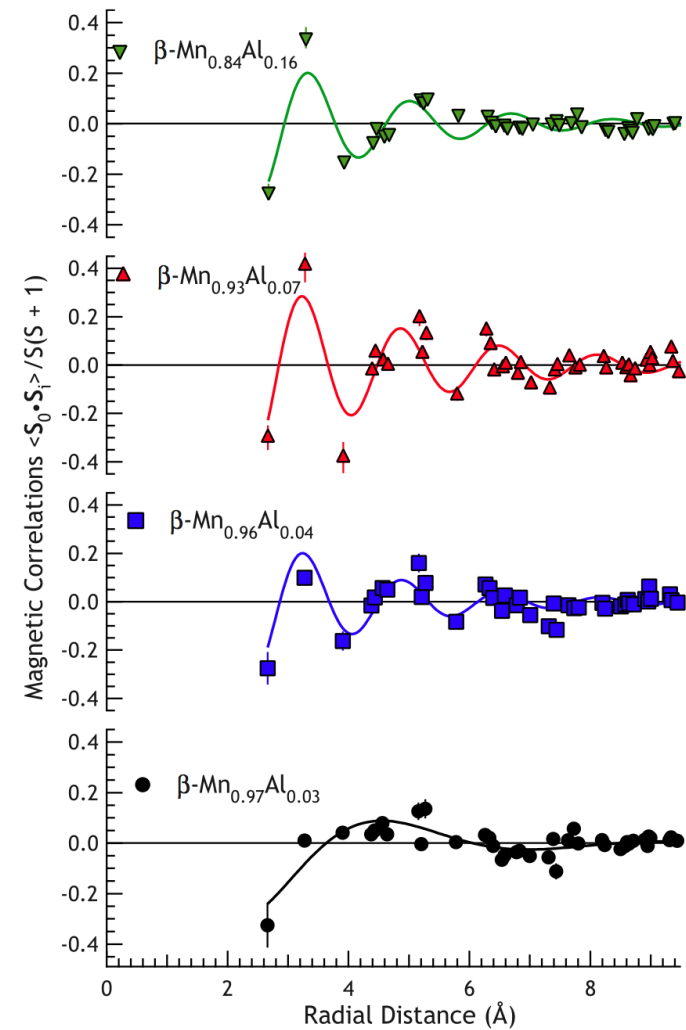
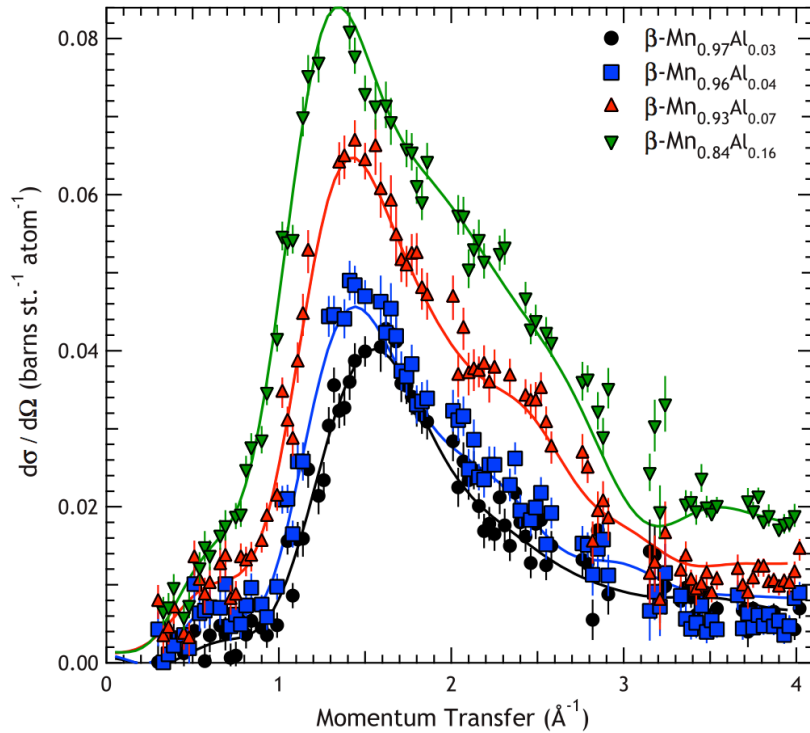
$$\mathbf{Q} \cdot \mathbf{r} = Qr \cos\theta$$

$$\int_{-\infty}^{\infty} f(\mathbf{Q}) \exp(i\mathbf{Q} \cdot \mathbf{r}) \cdot d\mathbf{r} = 4\pi \int_0^{\infty} r^2 f(r) \frac{\sin Qr}{Qr} \cdot dr$$

Short-range order in β -MnAl

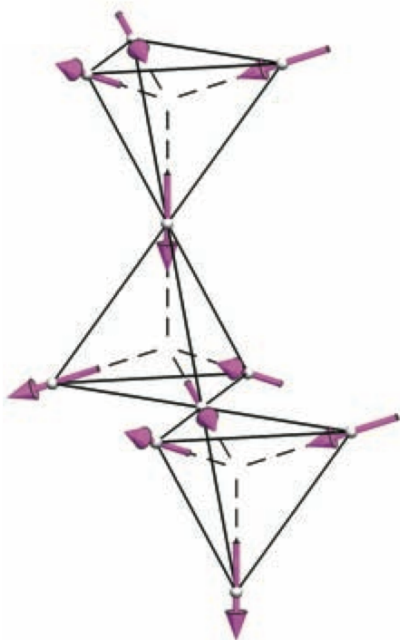
$$\left(\left| \langle \hat{V}^2 \rangle \right| - \left| \langle \hat{V} \rangle \right|^2 \right) \int z(\mathbf{r}) e^{i\mathbf{Q} \cdot \mathbf{r}} \cdot d\mathbf{r}$$

$$\propto S(S+1) + \sum_{R_n} \langle \mathbf{S}_0 \cdot \mathbf{S}_{R_n} \rangle \frac{\sin(QR_n)}{QR_n}$$

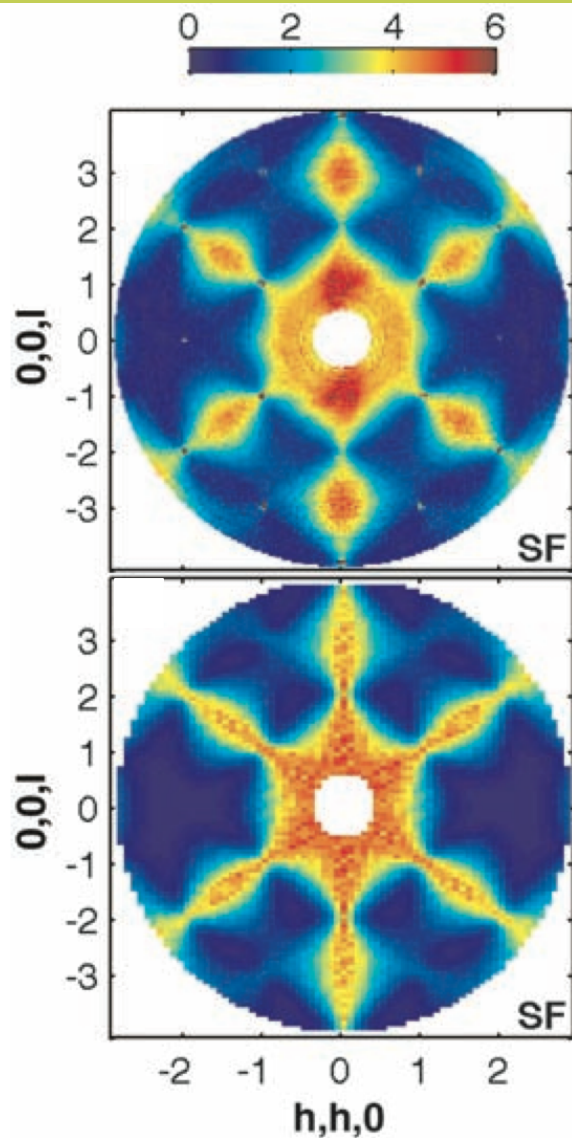


J. R. Stewart *et al.*, Phys. Rev. B **78** (2008) 014428

$\text{Ho}_2\text{Ti}_2\text{O}_7$ is a 'spin-ice' compound



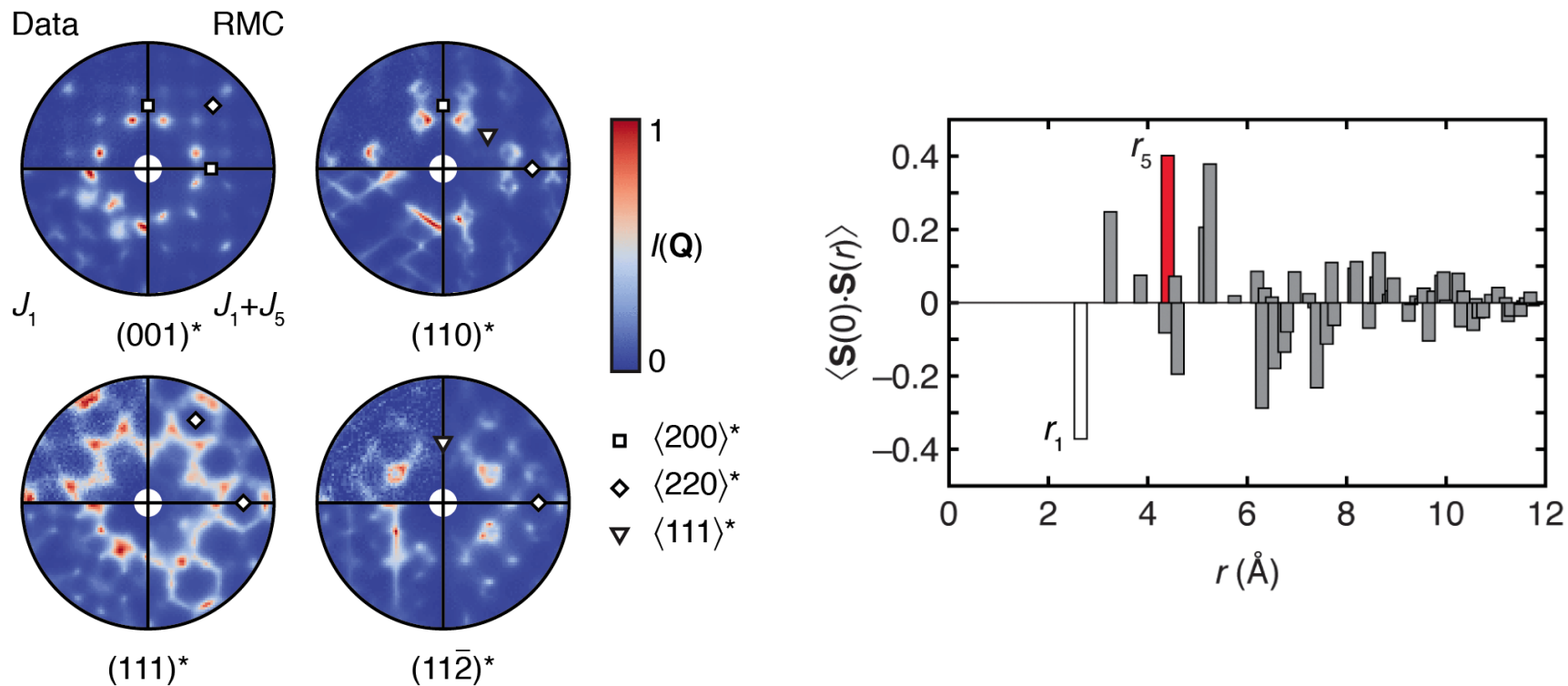
T. Fennell *et al.*, Science **326** (2009) 1177582



Data

Calculation

Use of “Reverse Monte-Carlo” to fit data



J. A. M. Paddison *et al.*, PRL **110** (2013) 267207

Conclusions:

- Learn your Fourier transforms
- Get used to using vectors
- Neutrons only ever see the components of the magnetization, $\mathbf{M}$, that are *perpendicular* to the scattering vector, $\mathbf{Q}$
- Magnetization is distributed in space, therefore the magnetic scattering has a *form factor*
- Neutron scattering can measure a susceptibility
- Be conscious of the relationship between time and energy, particularly for diffuse scattering
- Analysing diffuse scattering is non-trivial