

Outline

<u>This presentation might differ from the one which really</u> will be presented



Basics
Magnetic scattering
Spin manipulation
Instruments



Basics

P. 272 55

Una





Incoherent scattering

Reminder: The scattering cross section:

 $\frac{d\sigma}{d\Omega} = \sum_{i,j} b_i b_j \exp\left(i\vec{q}\left(\vec{R}_i - \vec{R}_j\right)\right)$

Suppose that at position R_i we can have different scattering length with a certain probability distribution

$$\langle b_i b_j \rangle = \langle b_i \rangle \langle b_j \rangle + \delta_{ij} \left(\langle b_i^2 \rangle - \langle b_i \rangle^2 \right)$$

$$\frac{d\sigma}{d\Omega} = \sum_{i,j} \langle b_i \rangle \langle b_j \rangle \exp\left(i\vec{q} \left(\vec{R}_i - \vec{R}_j\right)\right) + \sum_i \left(\langle b_i^2 \rangle - \langle b_i \rangle^2 \right)$$

 $I_{coherent}(\vec{q}) + I_{incoherent}(\vec{q})$



Isotop and spin incoherence

bi can have a distribution because:

- different isotopes exist

 the nucleus has a spin I => with the neutron 1/2 spin it forms two possible states

I+1/2 => 2(I+1/2)+1 states with scattering length b_+ I-1/2 => 2(I-1/2)+1 states with scattering length b_-

Isotope Incoherence

Spin Incoherence

Coherent

 $\langle b_i \rangle = b_+ \frac{I+1}{2I+1} + b_- \frac{I}{2I+1}$

Incoherent

$$\left(\langle b_i^2 \rangle - \langle b_i \rangle^2\right) = (b_+ - b_-)^2 \frac{I(I+1)}{(2I+1)^2}$$

the nuclear spin is usually randomly oriented EXCEPT very low T or very high B

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Let's define the polarization of the beam as: $ec{P}=2\left<ec{s}\right>=\left<ec{\sigma}\right>$ In terms of Pauli matrices

 $\sigma_x = \begin{pmatrix} 01\\10 \end{pmatrix} \qquad \sigma_y = \begin{pmatrix} 0-i\\i \ 0 \end{pmatrix} \qquad \sigma_z = \begin{pmatrix} 1 \ 0\\0-1 \end{pmatrix}$

Scattering length on a nucleus with spin I:

 $b = A + \frac{1}{2}B\hat{\sigma}\hat{I}$ $A = b_{+}\frac{I+1}{2I+1} + b_{-}\frac{I}{2I+1}$ $B = \frac{2(b_{+}-b_{-})}{2I+1}$

 σ_x and σ_y changes the spin state of the neutron => spin flip scattering σ_z does not => non spin flip scattering If the nuclear spin I is randomly oriented in space each one has 1/3 probability thus: Spin Incoherent scattering 2/3 spin flip 1/3 non spin flip P => -1/3P



Magnetic scattering

Magnetic scattering



Magnetic Interactions 1

The neutron has a 1/2 spin => magnetic moment $\mu_n = \gamma_n \mu_N$ μ_N nuclear Bohr magneton and $\gamma_n = -1.913$

With a magnetic field the interaction potential Lovesey 1986:

 $V(\vec{R}) = -\vec{\mu}_{n}\vec{B} = \gamma_{n}\mu_{N}\left[2\mu_{B}curl\left(\frac{\vec{s}\times\vec{R}}{|R^{3}|}\right) - \frac{e}{2m_{e}c}\left(\vec{p}_{e}\frac{\vec{\sigma}\times\vec{R}}{|R^{3}|} + \frac{\vec{\sigma}\times\vec{R}}{|R^{3}|}\vec{p}_{e}\right)\right]$ $\vec{s} = eletron spin operator, \vec{p}_{e} = eletron momentum operator$ $\vec{\sigma} = neutron spin operator$



Magnetic Interactions 2

The matrix element (scattering probability) becomes:

$$\langle k' | V_M | k \rangle = -r_0 \hat{\sigma} \hat{Q}_\perp \text{ with } r_0 = \frac{\gamma_n e^2}{m_e c^2}$$

$$\hat{Q}_\perp = \sum_i \exp\left(i\vec{q}\vec{r}_i\right) \left(\tilde{q} \times (\vec{s} \times \tilde{q}) - \frac{i}{\hbar |\vec{q}|} (\tilde{q} \times \vec{p}_i) \right) \text{ with } \tilde{q} = \frac{\vec{q}}{|\vec{q}|}$$

or in terms of the magnetization and changing to integral to account for the spatial extent of the electrons

$$\vec{Q} = -\frac{1}{2\mu_B} \int d\vec{r} \exp\left(i\vec{k}\vec{r}\right) \vec{M}\left(\vec{r}\right) \text{ and } \vec{Q}_{\perp} = \vec{Q} - \tilde{q}\left(\vec{Q}\tilde{q}\right)$$

Of the sample spins (magnetization) ONLY THE COMPONENT PERPENDICULAR to q contributes !!This is fundamentally different from the nuclear spin!



The scattered intensity is the Fourier transform of the self correlation function of the scattering length density

Alternatively if the Fourier transform of the scattering length density is F(q) $S(q) = F(q)F^*(q)$

If $F(q) = F_N(q) + F_M(q)$ in the most generic case there will be four terms: • nuclear

- magnetic
- nuclear magnetic interference
- chiral

ALC: NO



Unpolarized neutrons - comparable intensity to nuclear - identified by a priory knowledge - temperature dependence

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D20



3D Polanal X

50% spin flip scattering

neutron beam polarization direction



50% non spin flip scattering

neutron k_{out}

Q scattering vector

neutron kin



3D Polanal Y

spin flip scattering

neutron beam polarization direction



spin flip scattering

neutron k_{out}

Q scattering vector

neutron k_{in}



3D Polanal Z

50% non spin flip scattering

neutron beam polarization direction



50% spin flip scattering

neutron kout

Q scattering vector

neutron kin



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Polarized Neutrons

3D pol anal coils





3D pol anal simple

 $UP_X = N + \frac{1}{2}M + \frac{1}{3}I$ $DOWN_X = \frac{1}{2}M + \frac{2}{3}I$ $UP_Z = N + \frac{1}{2}M + \frac{1}{3}I$ $DOWN_Z = \frac{1}{2}M + \frac{2}{3}I$ $UP_Y = N + \frac{1}{3}I$ $DOWN_Y = M + \frac{2}{3}I$

N the nuclear scattering M the magnetic scattering I the incoherent scattering



Inelastc corrections

When the scattering is non negligibly quasielastic

50% spin flip scattering

neutron beam polarization direction



not quite ! 50% non spin flip scattering

neutron k_{out}

with distribution

Q scattering vector

neutron k_{in}



3D pol anal complicated

$$UP_{X} = N\frac{1+fg}{2} + \frac{1}{2}M - \frac{gf}{2}M\sin^{2}\varepsilon + \frac{3-gf}{6}I$$

$$DOWN_{X} = N\frac{1-g}{2} + \frac{1}{2}M + \frac{g}{2}M\sin^{2}\varepsilon + \frac{3+g}{6}I$$

$$g \text{ is the polarizing efficiency}$$

$$UP_{Z} = N\frac{1+fg}{2} + \frac{1}{2}M + \frac{3-gf}{6}I$$

$$DOWN_{Z} = N\frac{1-g}{2} + \frac{1}{2}M + \frac{3+g}{6}I$$

$$UP_{Y} = N\frac{1+fg}{2} + \frac{1-gf}{2}M + \frac{gf}{6}M\sin^{2}\varepsilon + \frac{3-gf}{6}I$$

$$UP_{Y} = N\frac{1+fg}{2} + \frac{1-gf}{2}M + \frac{gf}{2}M\sin^{2}\varepsilon + \frac{3-gf}{6}I$$



Spin manipulation

Spin manipulation



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Polarized Neutrons

Larmor precession

Quantum mechanical description of the neutron spin state:

 $|\chi
angle=a\,|+
angle+b\,|angle\,\,$ where a and b can be complex and $\sqrt{a^2+b^2}=1$ and conveniently $|\chi\rangle = e^{-i\varphi/2}\cos\frac{\theta}{2}|+\rangle + e^{i\varphi/2}\sin\frac{\theta}{2}|-\rangle$ this leads to $\langle \hat{\sigma}_x \rangle = \sin \theta \cos \phi, \ \langle \hat{\sigma}_y \rangle = \sin \theta \sin \phi, \ \langle \hat{\sigma}_x \rangle = \cos \theta$ Time evolution of the neutron spin in magnetic field $\frac{\partial \hat{\sigma}}{\partial t} = \frac{1}{\hbar} [\hat{\sigma}, H_s] = -\frac{\gamma}{2} \left[\hat{\sigma}, \left(\hat{\sigma} \vec{H} \right) \right] = \dots = -\gamma \left(\vec{H} \times \hat{\sigma} \right)$ The Polarization of the beam $P = <\sigma >$ behaves as a "classical" Larmor precession

 $\frac{\Delta \Phi}{\Delta x} \left[\frac{\deg}{cm} \right] = 2.65 \lambda \left[\mathring{A} \right] H \left[Gauss \right]$

 $\gamma_{\rm I} = 2957 \, \text{Hz/Gauss}$



Adiabaticity

What happens if the direction of the magnetic field changes in space?

The moving neutron will see a B field which changes its direction in time

Two limiting cases

if $\omega_{\rm B} \ll \omega_{\rm L}$ it will follow adiabatically

if $w_B \gg w_L$ it will start to precess around the new field direction

adiabatic example:

magnetic field direction

neutron trajectory



Mezei flipper

Non adiabatic example : Mezei flipper



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Polarized Neutrons

Real flippers





RF flipper



Makes up a π flipper • if $B_0 \gamma_L = \omega_{RF}$ • and B_{RF} is just enough for a π turn during the flight time



Haussler

Haussler Xtal Cu₂MnAl (111):

- $F_M(q) = F_N(q)$ for one of the Bragg reflections
- · easy to saturate
- grow single crystal
- controlled mosaicity
- low $\Lambda/2$ contamination (or filter)

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Polarized Neutrons

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Multilayer

303





$$= \frac{4N^2d^4(f_1 - f_2)^2}{\pi^2 n^4}$$

$$\frac{\lambda}{\lambda} = \frac{2d^2(f_1 - f_2)}{\pi}$$

$$\frac{\lambda}{2d}$$

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Polarized Neutrons

Supermirror







He³ absorbs only neutrons with spin antiparallel to the nuclear spin



He3

1000



Instruments

Instruments

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Polarized Neutrons



D7

JUD





For elastic scattering:

$$\varphi_{tot} = \frac{\gamma B_1 l_1}{v_1} - \frac{\gamma B_2 l_2}{v_2} = 0$$

For omega energy exchange:

$$\varphi_{tot} = \frac{\hbar \gamma B l}{m v^3} \omega + o \left(\left(\frac{\omega}{1/2m v^2} \right)^2 \right)$$

The probability of omega energy exchange:

 $S(q, \boldsymbol{\omega})$

The final polarization: $\langle \cos \varphi \rangle = \frac{\int \cos(\frac{\hbar \gamma Bl}{mv^3}\omega)S(q,\omega)d\omega}{\int S(q,\omega)d\omega} = S(q,t)$

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Polarized Neutrons

resolution correction







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Polarized Neutrons

When NOT to use NSE

Say you have a weak but well defined excitation



 $S(q, \omega) = 0.95 \cdot \delta(\omega) + 0.025 \cdot (\delta(\omega - \omega_0) + \delta(\omega + \omega_0)) \qquad S(q, t) = 0.95 + 0.05 \cdot \cos(\omega_0)$

Bad signal to noise, $\Delta \lambda / \lambda =>$ no good for Spin Echo

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10% marked polymer chain(H) in deuterated matrix of the same polymer melt

at short time => Rouse dynamics 1/tau ~ q⁴ at longer times starts to feel the "tube" formed by the other chains (deGennes)

D. Richter, B. Ewen, B. Farago, et al., Physical Review Letters 62, 2140 (1989).



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P. Schleger, B. Farago, C. Lartigue, et al., Physical Review Letters 81, 124 (1998).

Reptation

Spin Ice

'spin ice' materials $Ho_2 Ti_2 O_2$, $Dy_2 Ti_2 O_7$ and $Ho_2 Sn_2 O_7$ the spin is equivalent to the H displacement vector in water ice Paramagnetic echo => Only magnetic scattering gives echo !

Polarized Neutrons

Single exponential thermally activated

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G. Ehlers , Cornelius , A L, Orendac , M, Kajnakova , M, Fennell , T, Bramwell , S T, Gardner , Journal of Physics Condensed Matter 15, L9 (2003).



Over

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It's over folks ! thanks...